



# Sentinel-1 Toolbox

## Interferometry Tutorial

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# Interferometry Tutorial

The goal of this tutorial is to provide novice and experienced remote sensing users with step-by-step instructions on interferometric processing with SAR products.

In this tutorial you will process a pair of Stripmap Sentinel-1 products using interferometry.

## *What is Interferometry?*

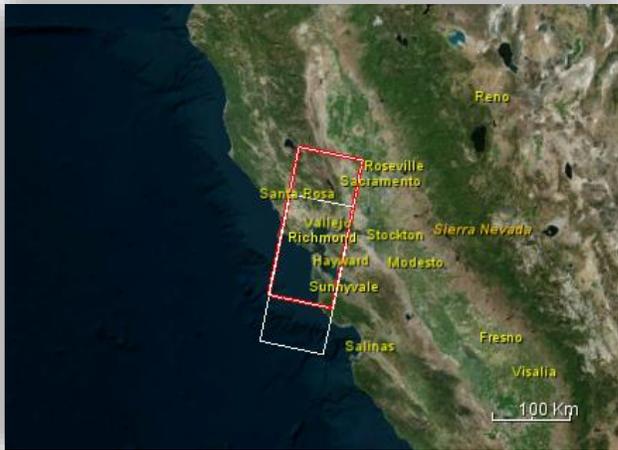
Interferometric synthetic aperture radar (InSAR) exploits the phase difference between two complex radar SAR observations taken from slightly different sensor positions and extracts information about the earth's surface.

A SAR signal contains amplitude and phase information. The amplitude is the strength of the radar response and the phase is the fraction of one complete sine wave cycle (a single SAR wavelength). The phase of the SAR image is determined primarily by the distance between the satellite antenna and the ground targets.

By combining the phase of these two images after coregistration, an interferogram can be generated whose phase is highly correlated to the terrain topography.

For an introduction to interferometric concepts, please see ESA's [InSAR Principles: Guidelines for SAR Interferometry Processing and Interpretation \(ESA TM-19\)](#).

## *Napa Valley Earthquake*



The 2014 South Napa earthquake occurred in and around the city of Napa, California on August 24 at 3:20 a.m. local time, measuring at 6.0 on the moment magnitude scale.

This is the first earthquake for which the surface deformation has been measured by ESA's Sentinel-1 satellite.

Two Stripmap SLC data products are available at the links below. One product is acquired on 7 August and another captured on 31 August.

[http://step.esa.int/auxdata/tutorials/s1tbx/NapaValley/S1A\\_S1\\_SLC\\_1SSV\\_20140807T142342\\_20140807T142411\\_001835\\_001BC1\\_05AA.SAFE.zip](http://step.esa.int/auxdata/tutorials/s1tbx/NapaValley/S1A_S1_SLC_1SSV_20140807T142342_20140807T142411_001835_001BC1_05AA.SAFE.zip)

[http://step.esa.int/auxdata/tutorials/s1tbx/NapaValley/S1A\\_S1\\_SLC\\_1SSV\\_20140831T142335\\_20140831T142403\\_002185\\_002356\\_C2E5.SAFE.zip](http://step.esa.int/auxdata/tutorials/s1tbx/NapaValley/S1A_S1_SLC_1SSV_20140831T142335_20140831T142403_002185_002356_C2E5.SAFE.zip)

## Opening a Pair of SLC Products

In order to perform interferometric processing, the input products should be two or more Single Look Complex (SLC) products over the same area acquired from slightly different satellite positions.

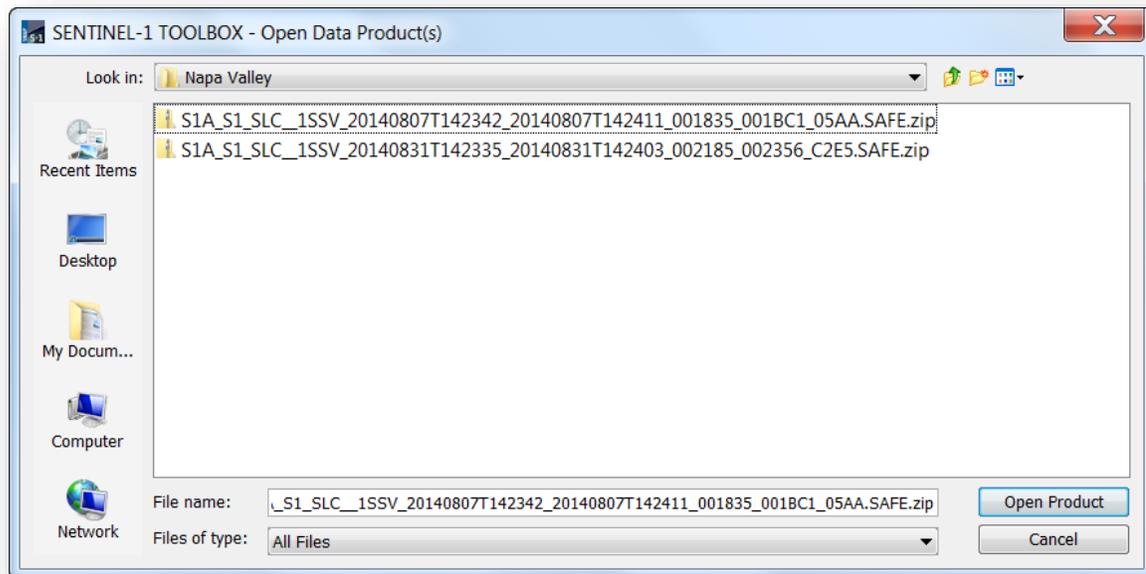
The toolbox can support SLC products for InSAR from:

- SENTINEL-1
- RADARSAT-2
- TerraSAR-X
- ENVISAT ASAR
- ERS 1&2
- Cosmo-Skymed
- ALOS 1&2

**Step 1 - Open the products:** Use the  **Open Product** button in the top toolbar and browse for the location of the products.

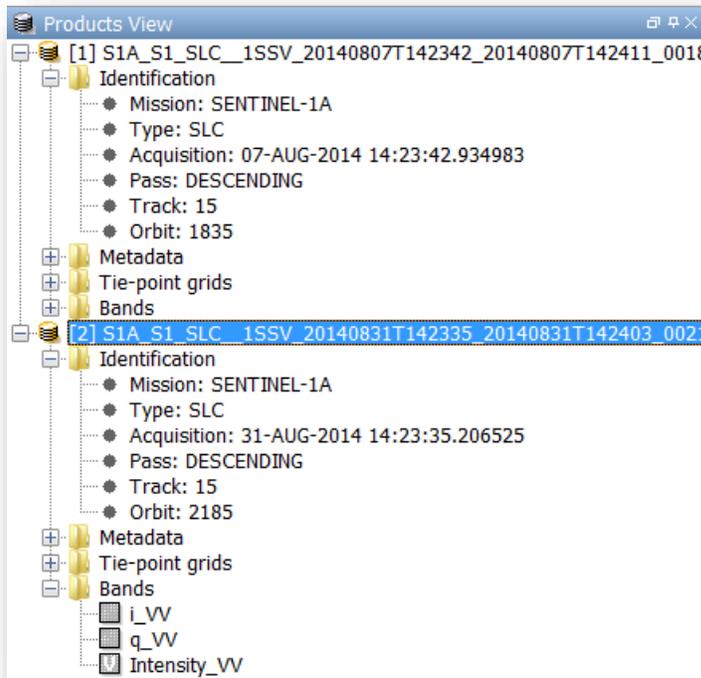
Select the **manifest.xml** file from each Sentinel-1 product folder and press **Open Product**.

If the products are zipped, you may also select the zip files.



### Opening a Product

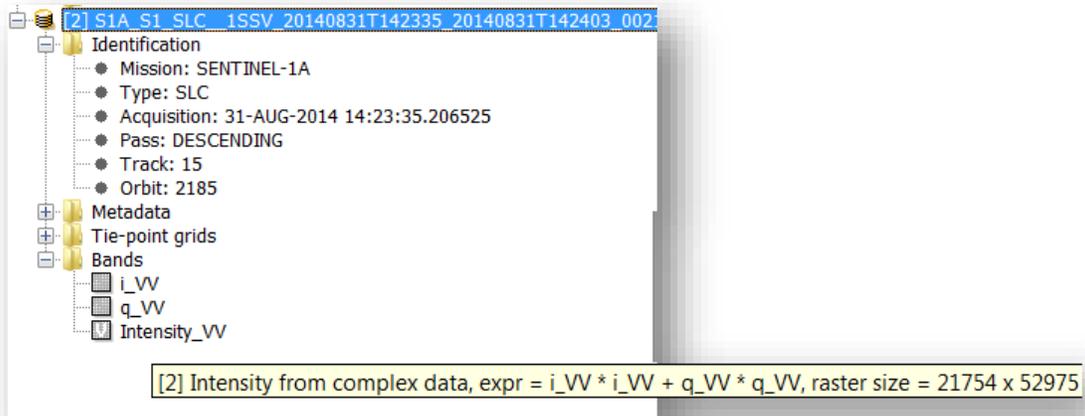
**Step 2 - View the product:** In the **Products View** you will see the opened products. Within the product bands, you will find two bands containing the real (i) and imaginary (q) parts of the complex data. The i and q bands are the bands that are actually in the product.



### Products View

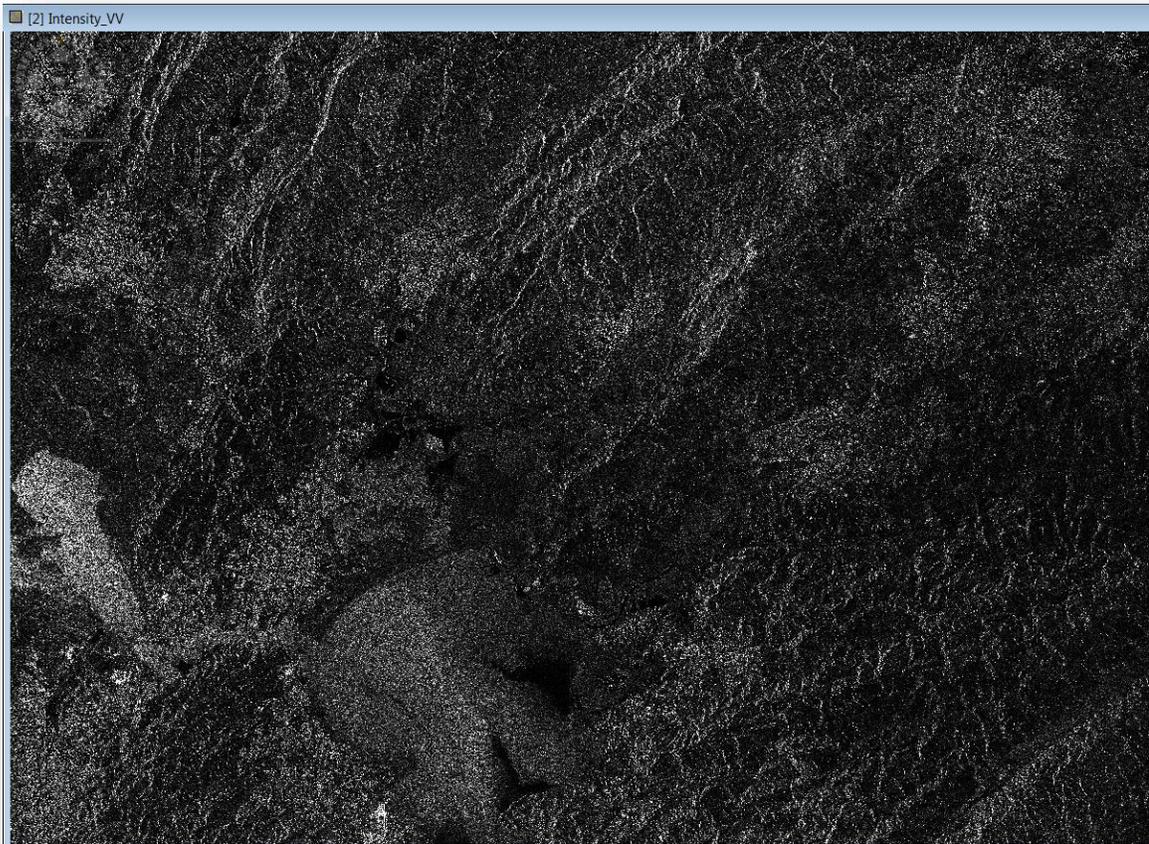
The virtual Intensity band is there to assist you in working with complex data.

Virtual bands are the result of creating a band using the Band Maths feature. If you hover the mouse over top of the Intensity virtual band, you will see the expression that it consists of. In this case, the Intensity is  $i \times i + q \times q$ .



### Intensity Band Tool-tip

**Step 3 - View a band:** To view the data, double-click on the **Intensity\_HH** band. Zoom in using the mouse wheel and pan by clicking and dragging the left mouse button.



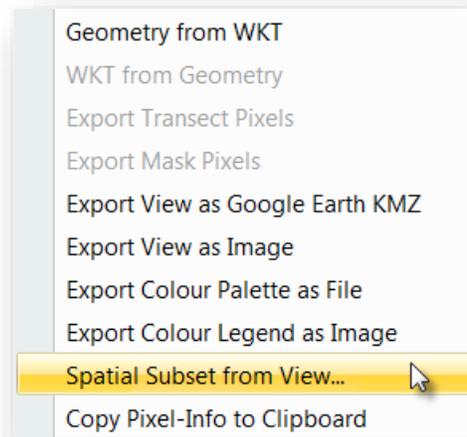
### Intensity Band

Pan and zoom to the hilly area in the middle of the scene.

### ***Creating a Subset***

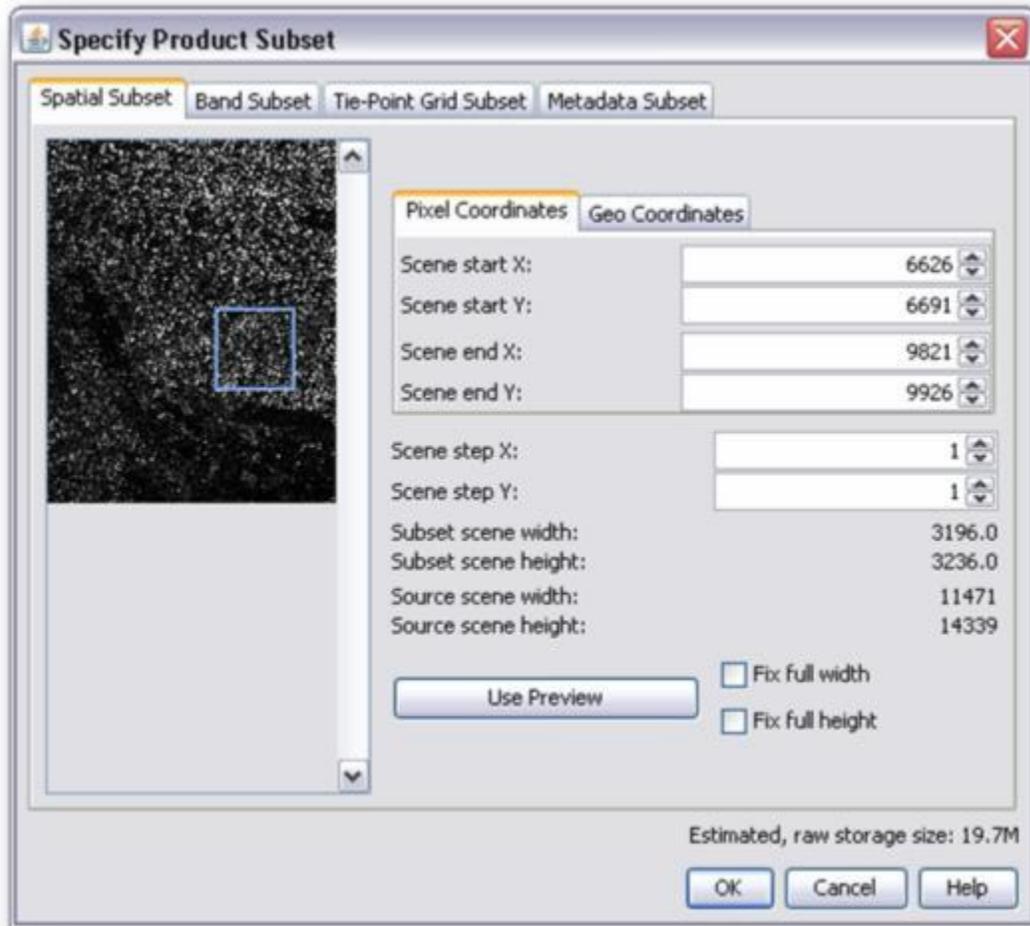
To reduce the amount of processing needed, you may create a subset around the particular area in which you are interested.

**Step 4 - Create a subset from the view:** Once you have zoomed and panned to your area of interest, right click and select **Spatial Subset from View** in the **context** menu.



**Select Spatial Subset from View**

The subset dialog will automatically select the area you were viewing.



### Specifying a Product Subset

By default all bands will be included in the subset. Press **OK** to create the subset.

For InSAR processing, you generally only need one polarization. If you have multiple polarizations in your input product, you may use the subset operator to select only co-polarized bands (HH or VV). Likewise, you could use the **BandSelect** operator to select all bands for a particular polarization.

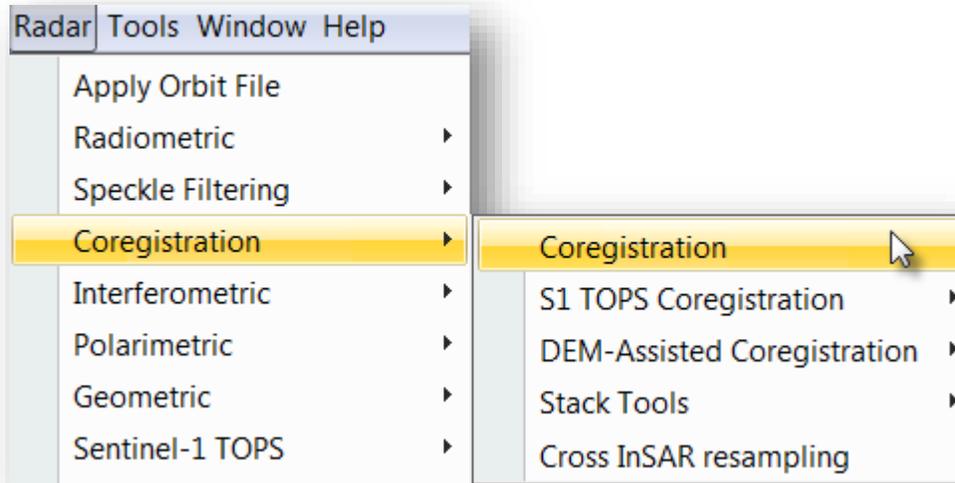
When the new subset product appears in the products view, press the save button  to save the product.

## Coregistering the Data

For interferometric processing, two or more images must be coregistered into a stack. One image is selected as the master and the other images are the slaves. The pixels in slave images will be moved to align with the master image to sub-pixel accuracy.

Coregistration ensures that each ground target contributes to the same (range, azimuth) pixel in both the master and the slave image.

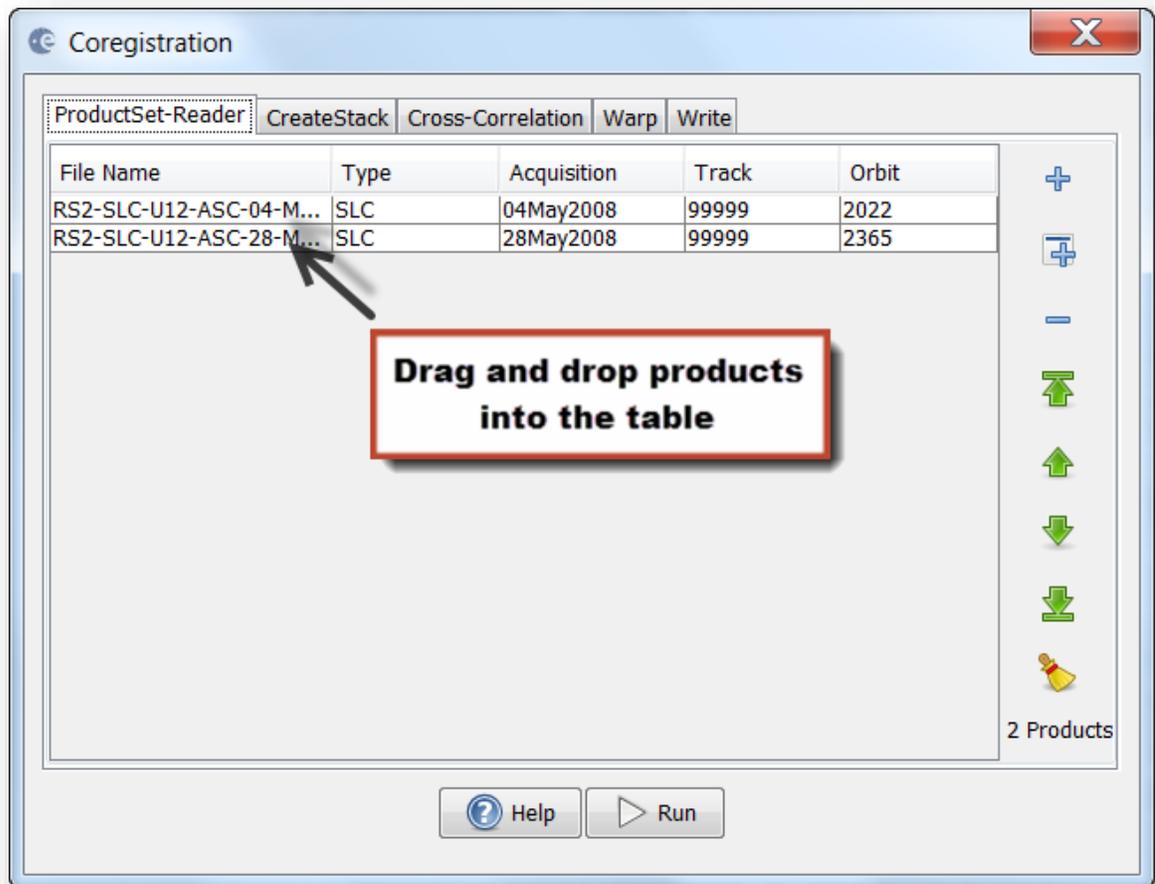
**Step 5 - Coregister the images into a stack:** Select **Coregistration** in the **Coregistration** menu.



### Select Coregistration

Drag and drop the products from the **Products View** to the table in the **Coregistration** dialog.

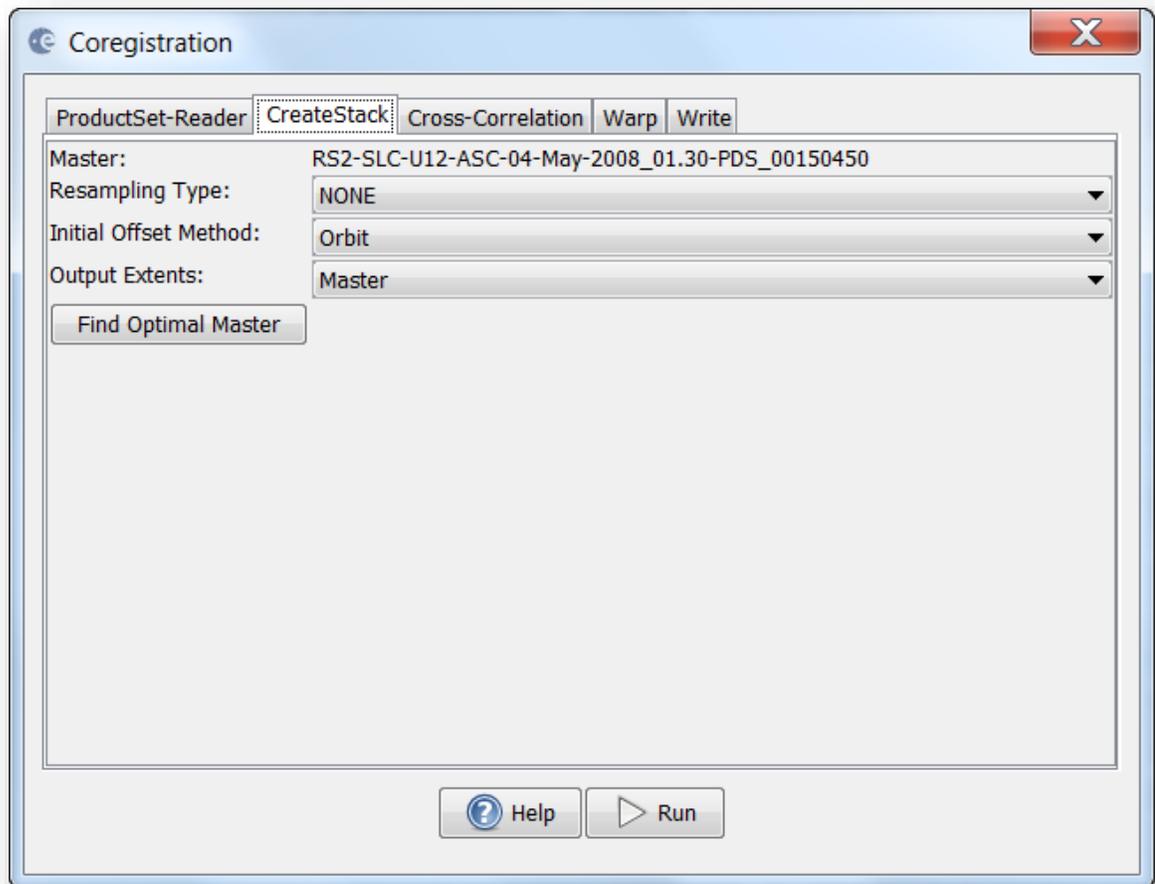
Drag and drop first the subset product. This will be your **master** image. Then drag and drop the other product. This will be your **slave** image.



### Add Products into the Coregistration Dialog

You could also press the **Add Opened** button to add all products currently open in the Products View.

In the **Create Stack** tab, the bands for the master image and the slave images should already be selected for you based on the order of the products given in the previous table.



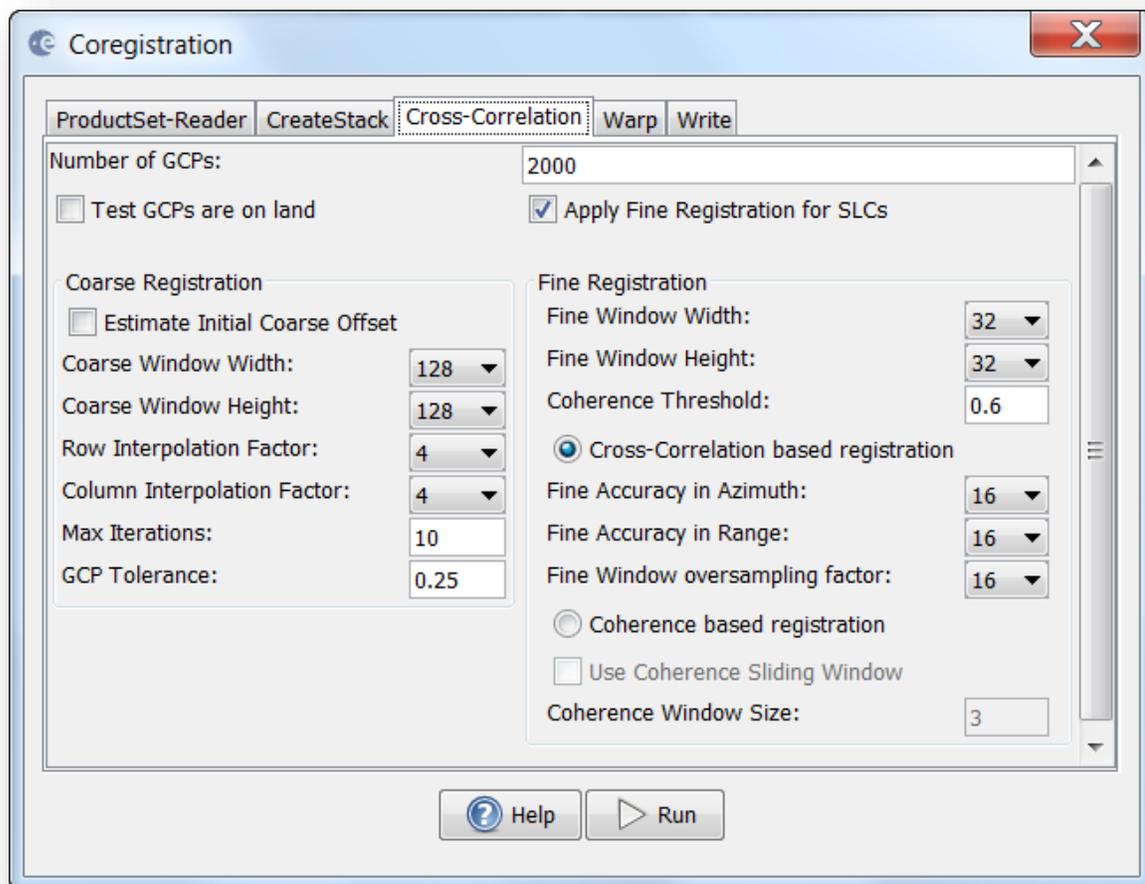
### Create Stack

For InSAR processing, it is best not to resample the images at this step. Select Resampling Type: **NONE** and Output Extents of the **Master**.

If you have several products to coregister, you may press **Find Optimal Master** to search for the best image to use as the master image based on perpendicular and temporal baselines.

In the **Cross-Correlation** tab, specify the number of Ground Control Points (GCPs) to use. The GCPs will be used as the center of a cross correlation window which will find the corresponding position from the slave image to the master image.

Trouble shooting: If after processing you receive an error such as "not enough GCPs survived", this means that the software failed to find a good correlation with the current parameters. Try increasing the number of GCPs to possibly 1000. If the offset between master and slave images is large, you will need to increase the coarse registration window dimensions to possibly 256 or 512.

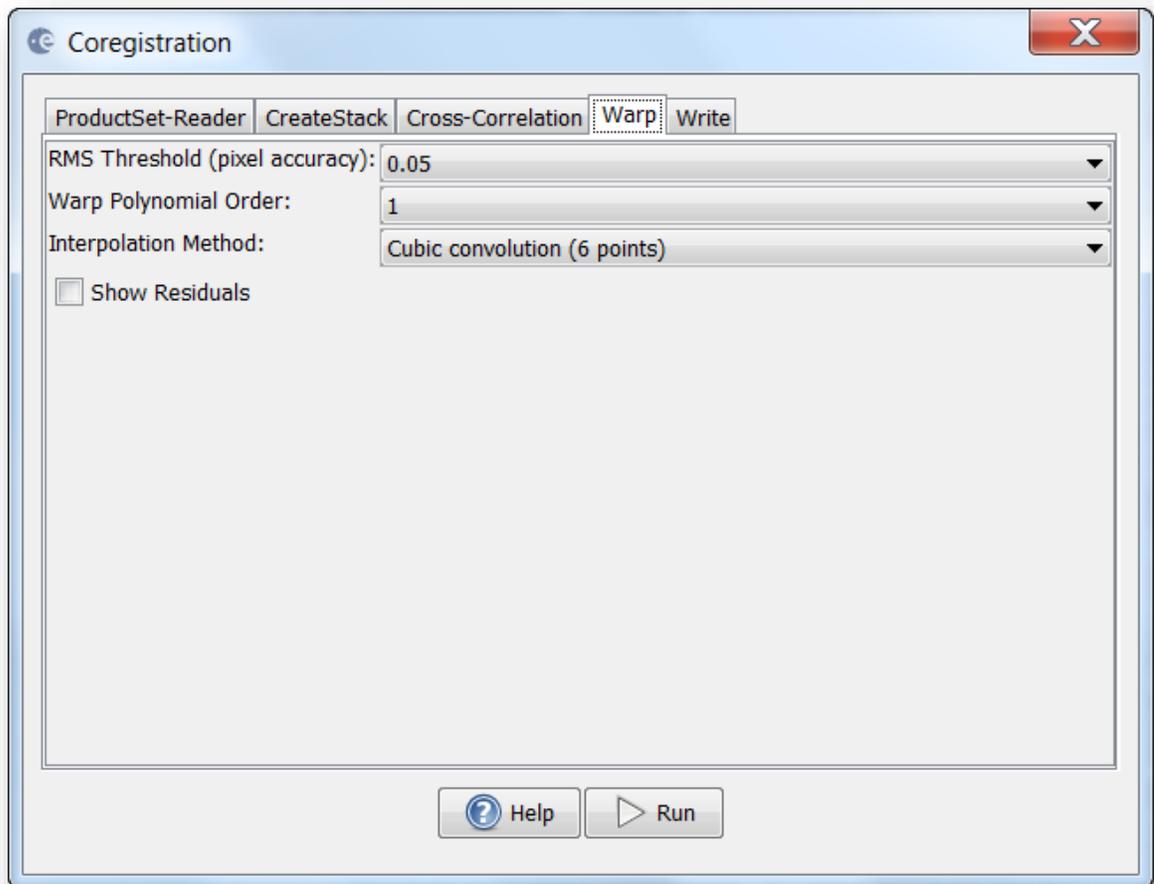


### Define the Correlation Windows

In the **Warp** tab, specify the significance level for outlier removal (RMS threshold) to a pixel accuracy of 0.01 and select the interpolation method.

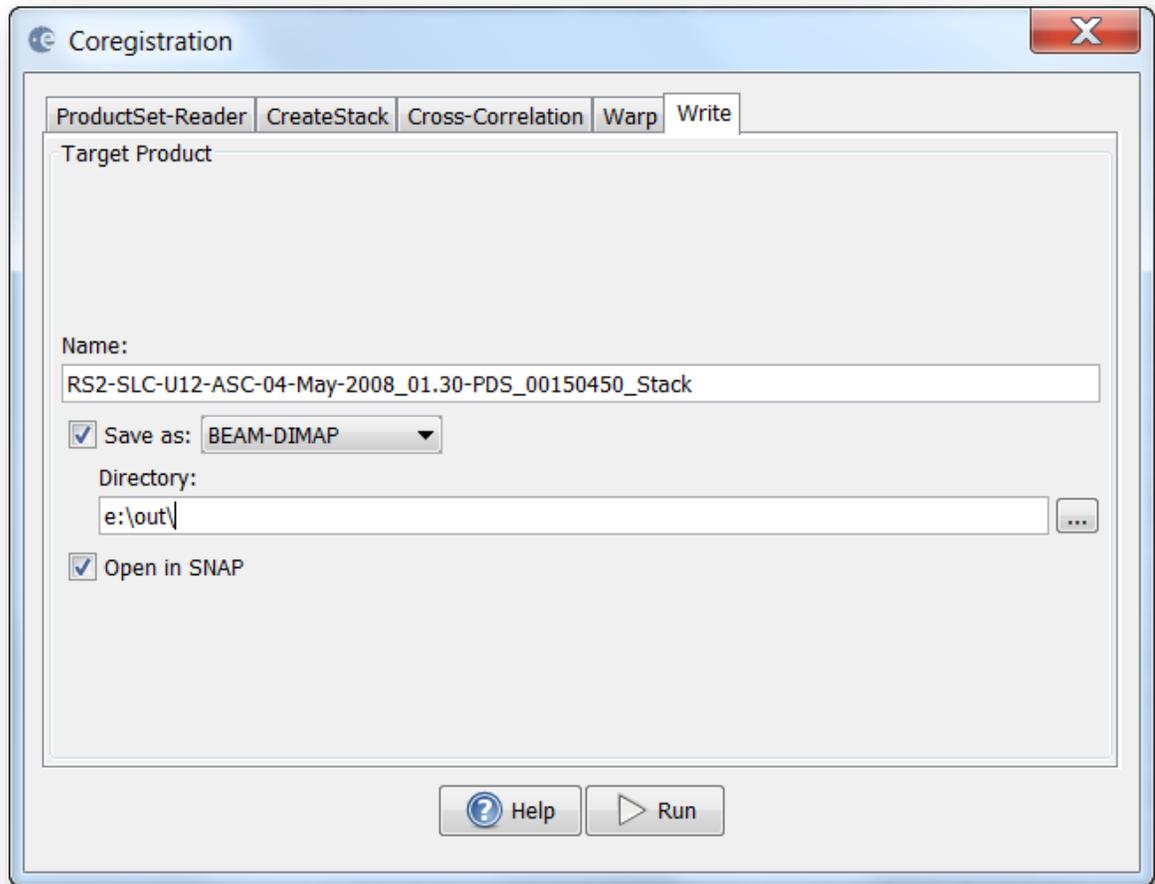
The lower you can get the RMS threshold while having enough GCPs surviving, the better the coregistration result will be. Coregistration accuracy plays a critical role in achieve good interferometric results.

The warp polynomial order applies a linear translation for order 1. Higher order warps should only be used when the images have been greatly distorted.



**Specify the Significance Level via the RMS Threshold**

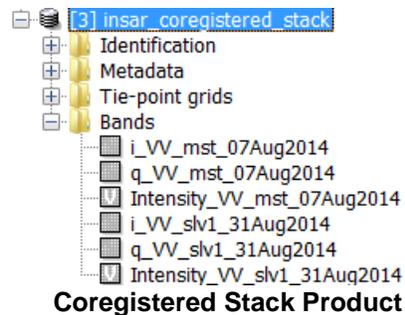
In the **Write** tab, specify the output folder and the target product name.



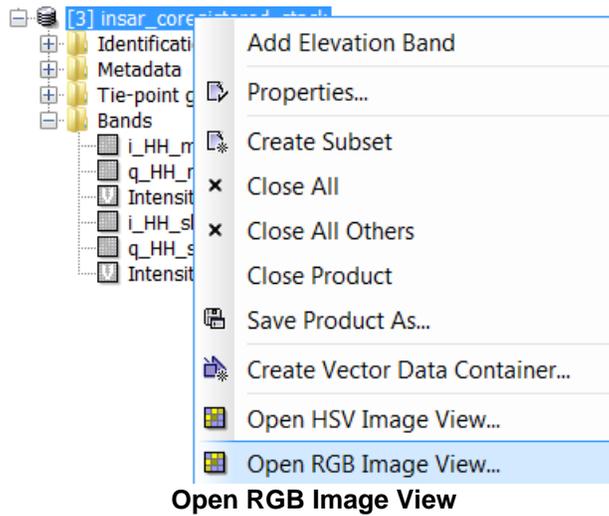
### Specify the output name, format and folder

Press **Process** to begin processing the coregistration.

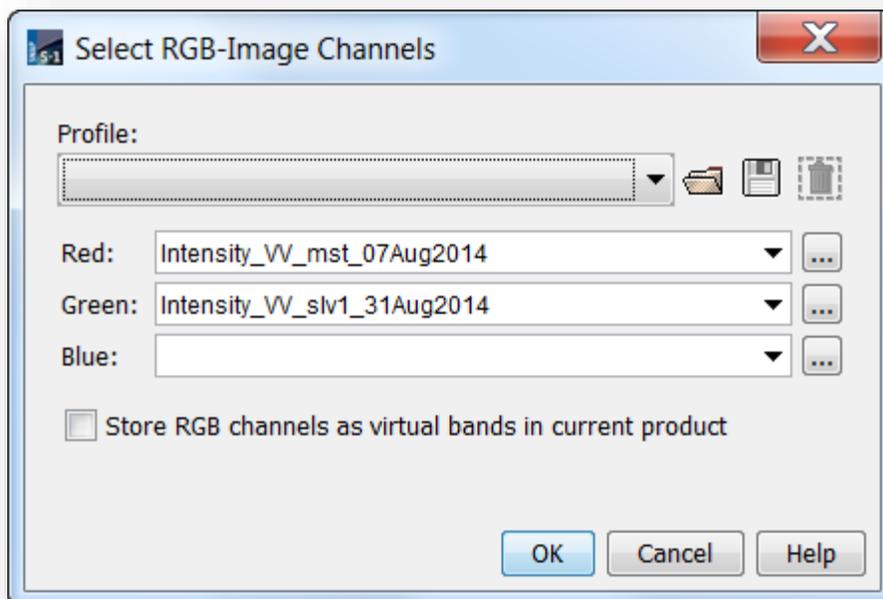
The resulting coregistered stack product will appear in the **Products View**.



To view the results of the coregistration in a colour view, right-click on the product name and select **Open RGB Image View**.

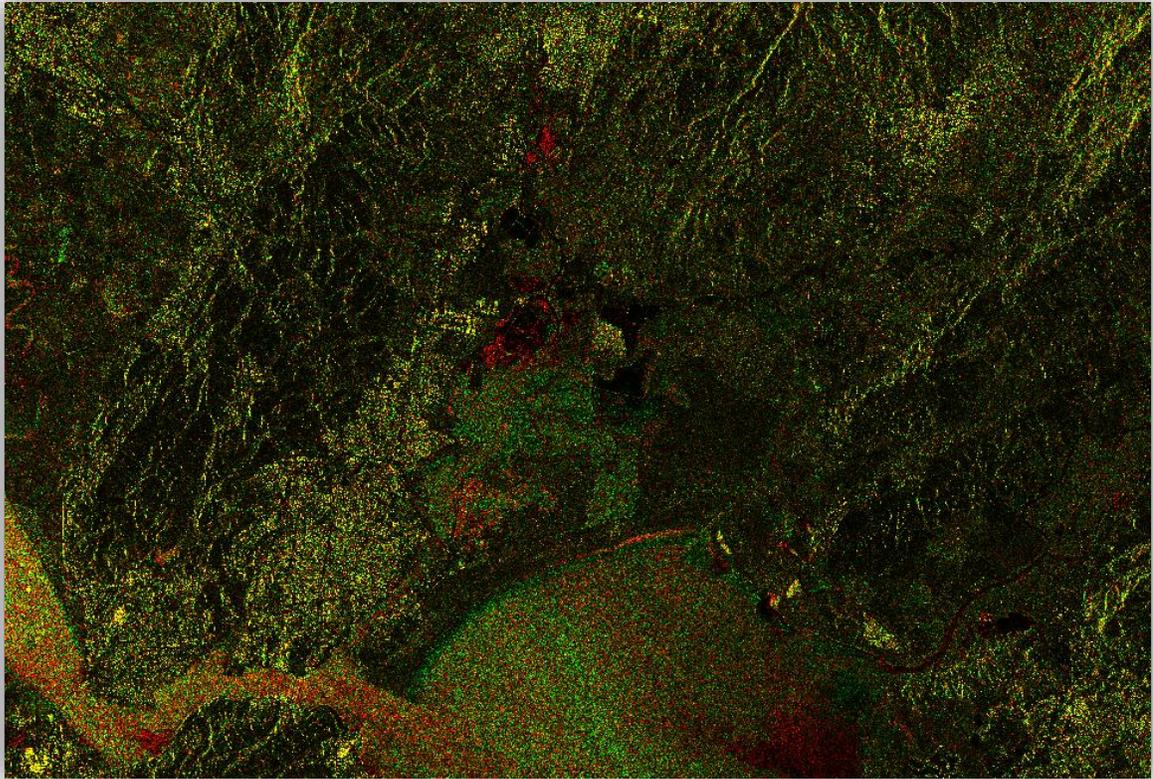


In the RGB channel selection dialog, the Intensity bands for both images will automatically be selected. Select the master image in red and slave image in green then press **OK**.



**Select master in red and slave in green**

The RGB View can be useful for amplitude change detection. In this image, you will see things that have changed in red or green and things that have not changed in yellow. It is also a visual indication that the coregistration has properly aligned both images. The resulting RGB view should look mostly yellow. Poor registrations will have badly lined up terrain.



RGB View

### ***What Interferometric Tools are Available?***

Once the images have been accurately coregistered, you can use the interferometric tools available in the Toolbox including:

- Range and Azimuth Spectral Filtering
- Coherence Estimation
- Interferogram Formation
- Topographic Phase Removal
- Phase Filtering
- Phase Unwrapping
- Unwrapped Phase to Height Conversion

### ***Spectral Filtering***

Spectral filtering improves the signal-to-noise ratio (SNR) in the interferogram. This noise reduction results from filtering out non overlapping parts of the spectrum. This spectral non overlap in range between master and slave is caused by a slightly different viewing angle of both sensors. The longer the perpendicular baseline is, the smaller the overlapping part.

The Range Filter operator filters the spectras of a stack of SLC images in the range direction. This Azimuth Filter operator filters the spectras of stack of SLC images in the azimuth direction. Spectral filtering is optional and may sometimes not affect the data very much.

## Interferogram Formation and Coherence Estimation

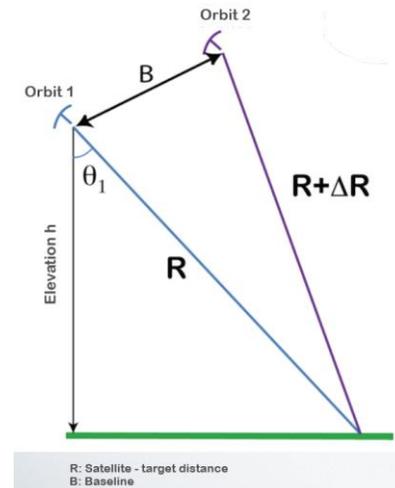
The interferogram is formed by cross multiplying the master image with the complex conjugate of the slave. The amplitude of both images is multiplied while the phase represents the phase difference between the two images.

The interferometric phase of each SAR image pixel would depend only on the difference in the travel paths from each of the two SARs to the considered resolution cell.

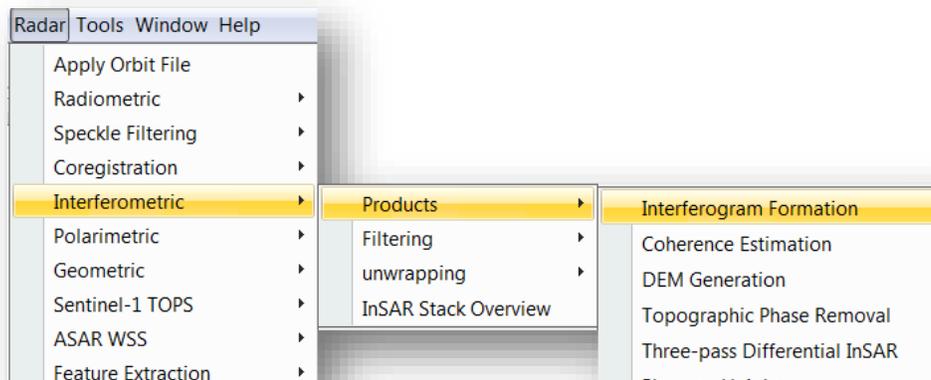
The interferometric phase variation  $\Delta\phi$  is then proportional to  $\Delta R$  divided by the transmitted wavelength  $\lambda$ .

$$\phi_1 = \frac{4\pi R}{\lambda}, \quad \phi_2 = \frac{4\pi(R + \Delta R)}{\lambda}$$

$$\Delta\phi = \phi_2 - \phi_1 = \frac{4\pi\Delta R}{\lambda}$$



**Step 6 - Form the Interferogram:** Select the stack and select **Interferogram Formation** from the **InSAR Products** menu.

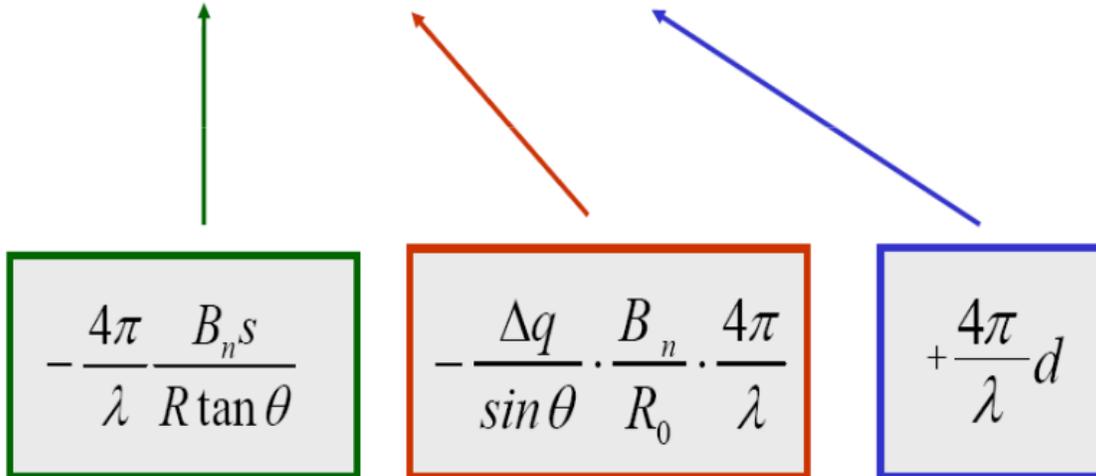


### Select Interferogram Formation

The phase difference can have contributions from five different sources:

- $\Delta\phi_{flat}$  is called flat Earth phase which is the phase contribution due to the earth curvature.
- $\Delta\phi_{elevation}$  is the topographic contribution to the interferometric phase.
- $\Delta\phi_{displacement}$  is the surface deformation contribution to the interferometric phase.

- $\Delta\phi_{atmosphere}$  is the atmospheric contribution to the interferometric phase. It is introduced due to the atmospheric humidity, temperature and pressure change between the two acquisitions.
- $\Delta\phi_{noise}$  is the phase noise introduced by temporal change of the scatterers, different look angle, and volume scattering.

$$\Delta\phi = \Delta\phi_{flat} + \Delta\phi_{elevation} + \Delta\phi_{displacement} + \Delta\phi_{atmosphere} + \Delta\phi_{noise}$$


$$-\frac{4\pi}{\lambda} \frac{B_n s}{R \tan \theta}$$

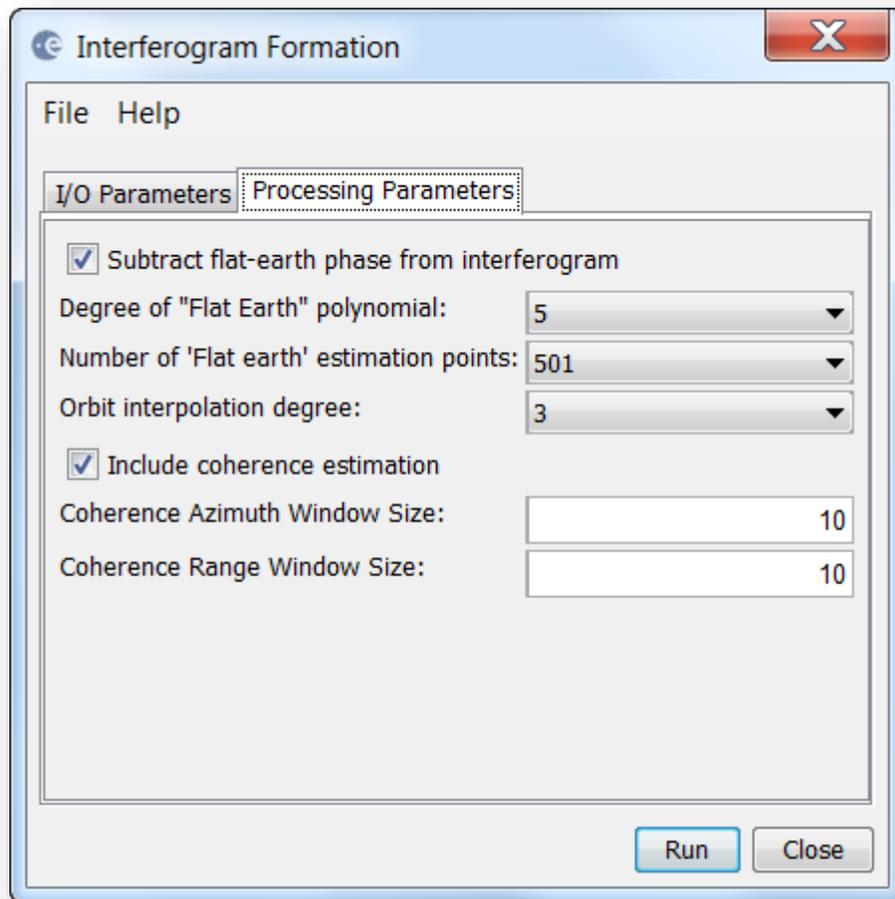
$$-\frac{\Delta q}{\sin \theta} \cdot \frac{B_n}{R_0} \cdot \frac{4\pi}{\lambda}$$

$$+\frac{4\pi}{\lambda} d$$

**Contributors to SAR Interferometric Phase**

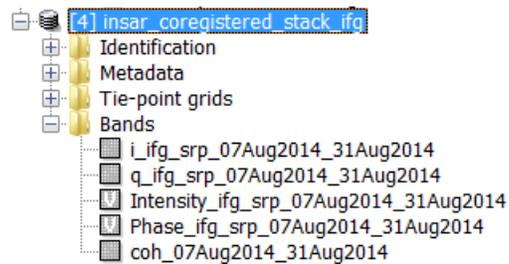
Through the interferometric processing, we shall try to eliminate other sources of error to be left with only the contributor of interest which is typically the elevation or the displacement.

In the interferogram formation step we shall remove the flat-Earth phase. The flat-Earth phase is the phase present in the interferometric signal due to the curvature of the reference surface. The flat-Earth phase is estimated using the orbital and metadata information and subtracted from the complex interferogram.

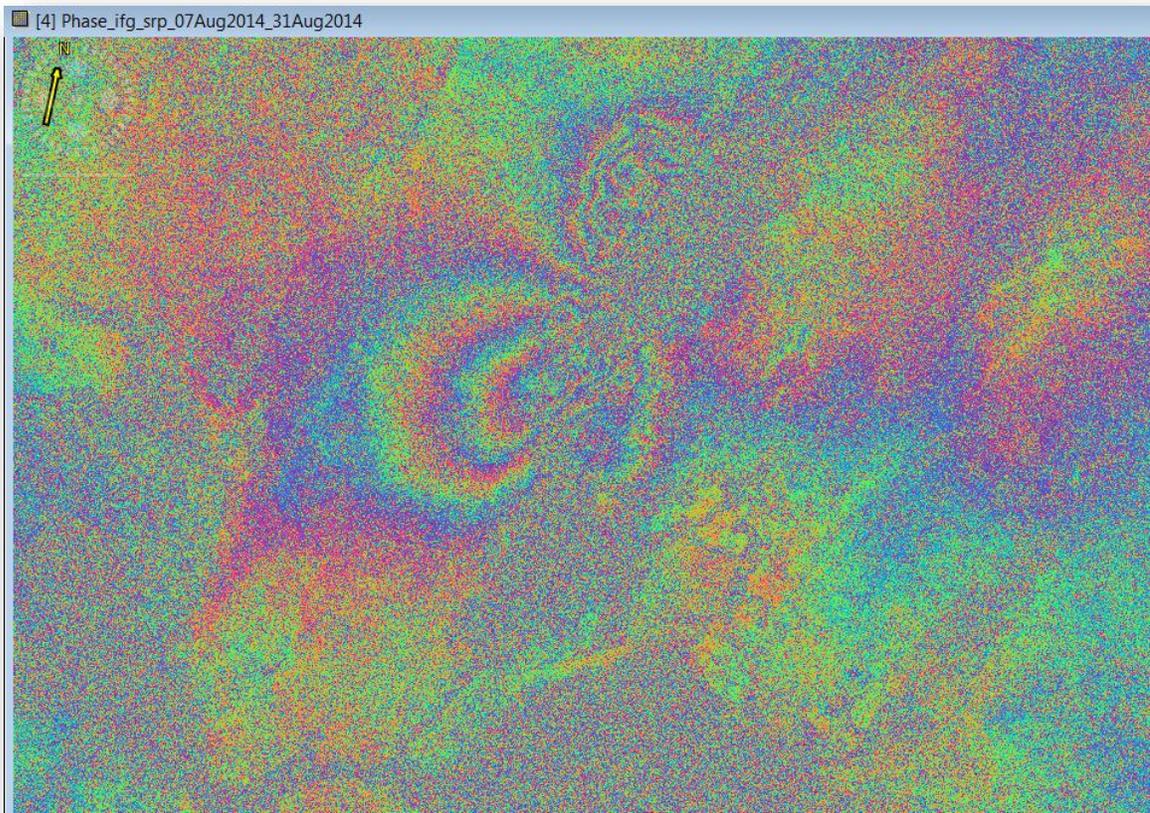


### Interferogram Dialog

The interferogram product produced will contain a band for the interferometric phase.



### Interferometric Phase Band



### Interferometric Phase Band

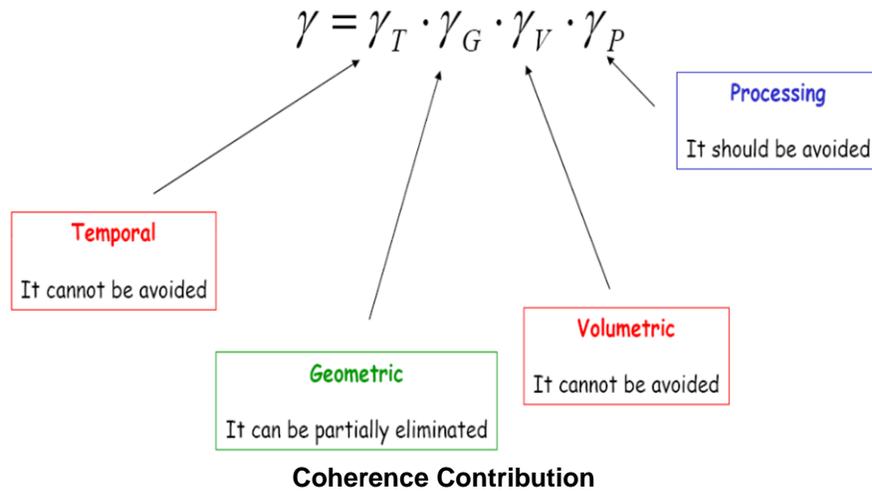
Interferometric fringes represent a full  $2\pi$  cycle. Fringes appear on an interferogram as cycles of arbitrary colours, with each cycle representing half the sensor's wavelength. Relative ground movement between two points can be calculated by counting the fringes and multiplying by half of the wavelength. The closer the fringes are together, the greater the strain on the ground.

Flat terrain should produce a series of regularly spaced, parallel fringes. Any deviation from a parallel fringe pattern can be interpreted as topographic variation.

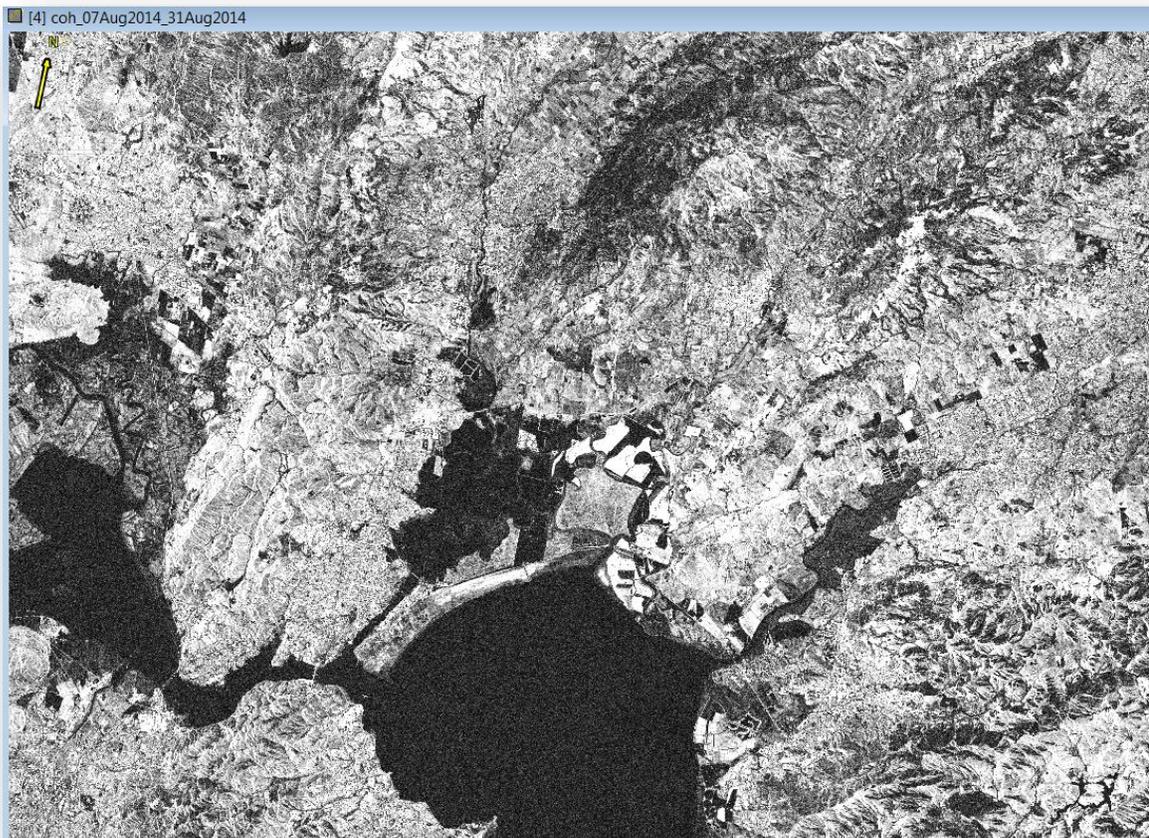
With the same operator, you may also generate the coherence estimation in the same processing step as the interferogram.

The coherence between master and slave images can show you if the images have strong similarities and are therefore good candidates for generating a DEM. Loss of coherence can produce poor interferometric results.

Loss of coherence could be caused by temporal (time between acquisitions), geometric (orbit errors), volumetric (vegetation) or processing.



The coherence band shows how similar each pixel is between the slave and master images in a scale from 0 to 1. Areas of high coherence will appear bright. Areas with poor coherence will be dark. In the image, vegetation is shown as having poor coherence and buildings have very high coherence.

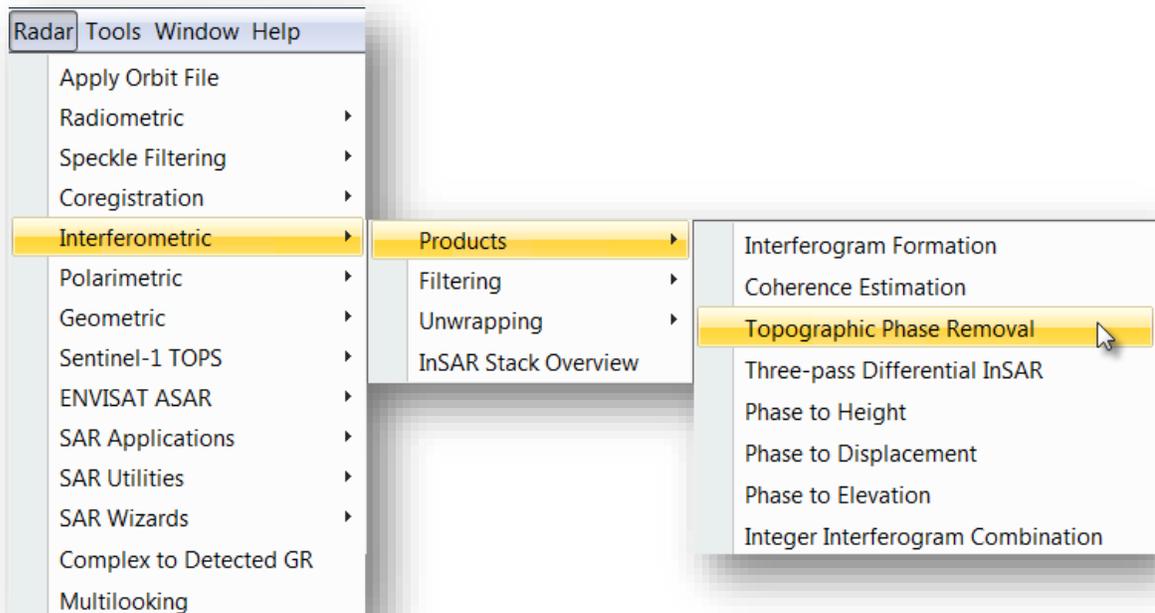


**Coherence Band**

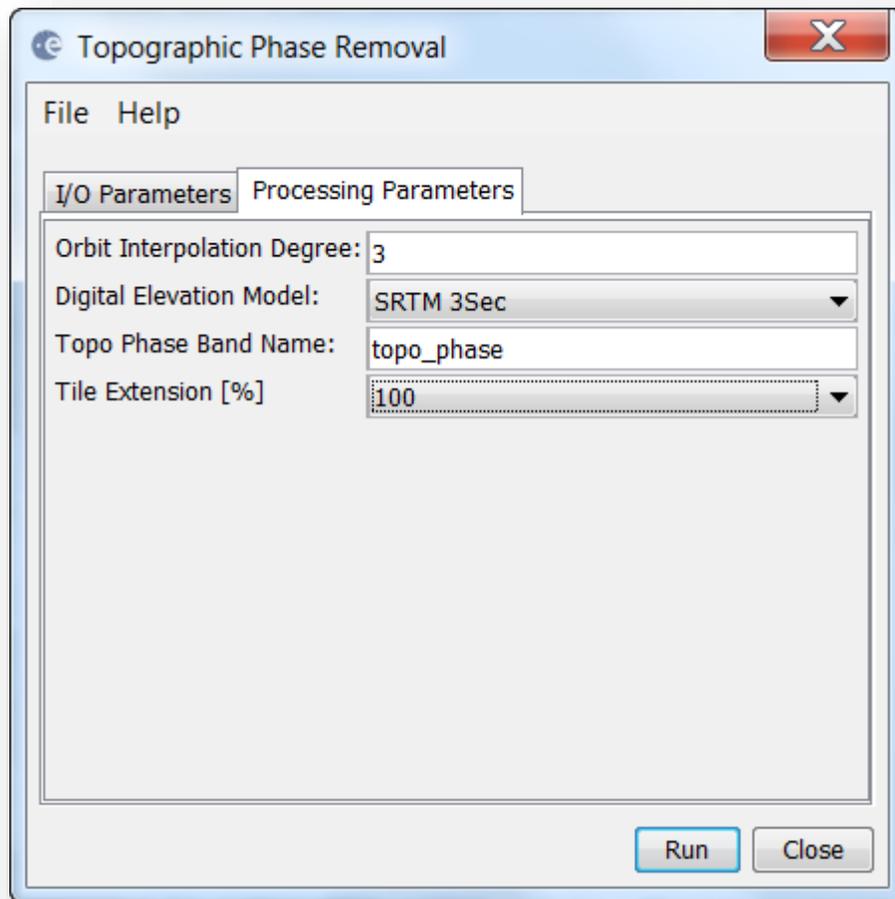
## ***Topographic Phase Removal***

The Interferogram can then be flattened by removing the topographic phase. The operator will simulate an interferogram based on a reference DEM and subtract it from the processed interferogram.

**Step 7 - Remove Topographic Phase:** Select the Interferogram product and go to the **Interferometric Products** menu. Select **Topographic Phase Removal**.

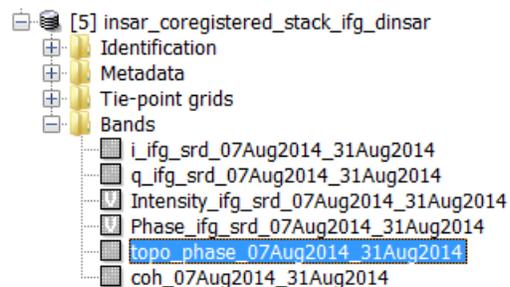


**Select Topographic Phase Removal**

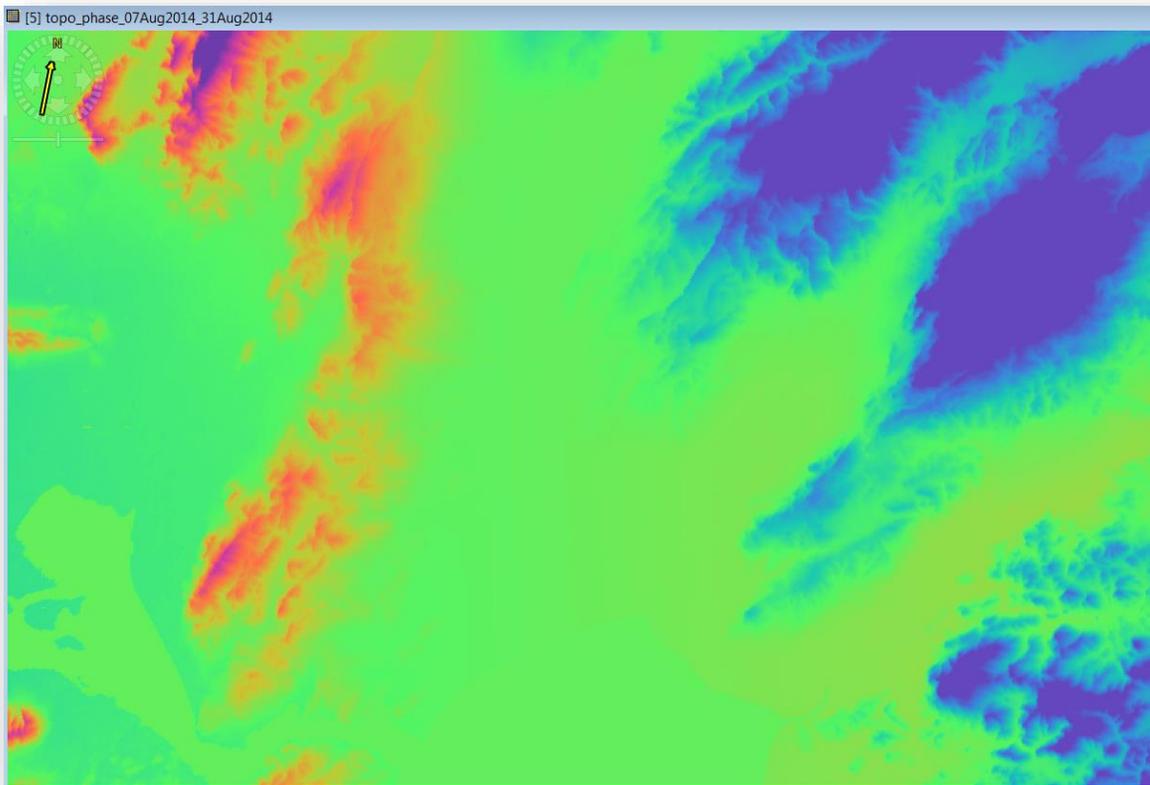


### Topographic Phase Removal Dialog

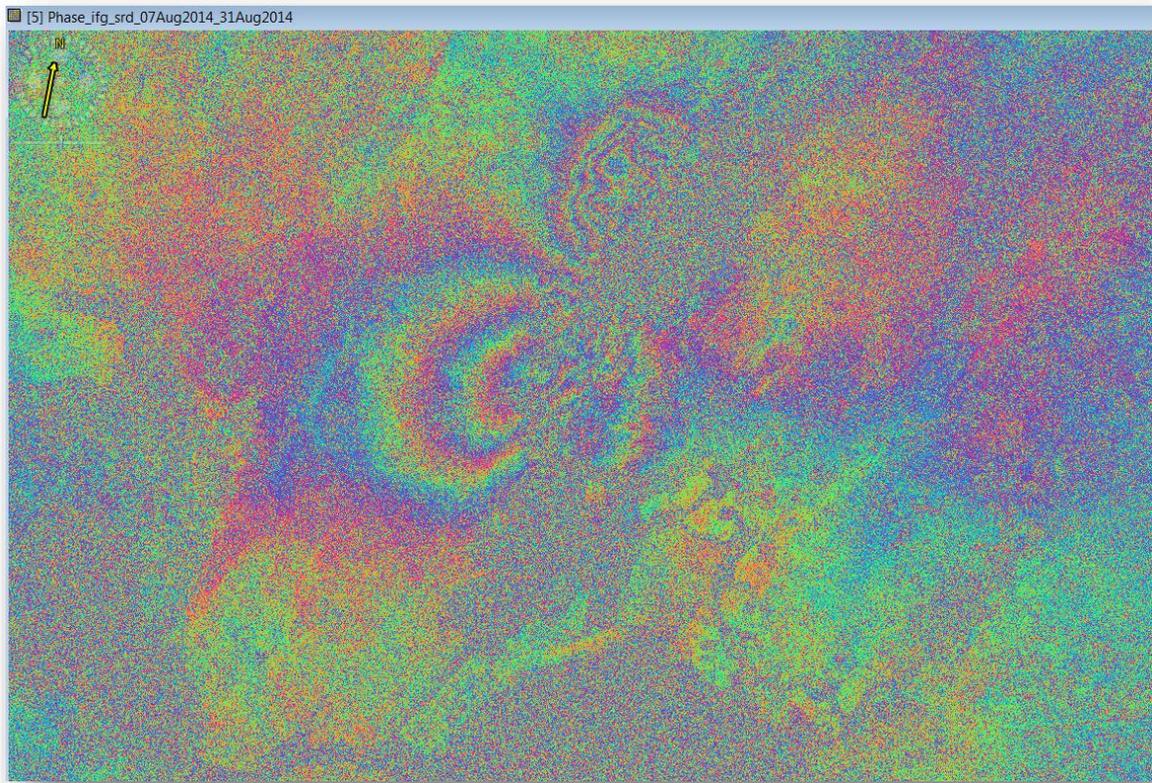
The resulting product will have an interferogram with topographic phase removed and a band for the topographic phase.



### Topographic Phase



**Topographic Phase Band**



**Topographic Phase Removed**

### ***Phase Filtering***

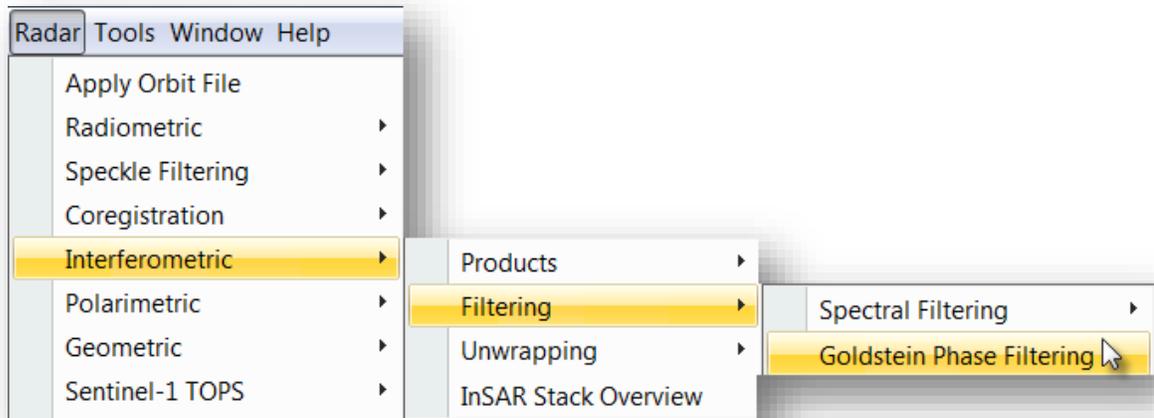
Interferometric phase can be corrupted by noise from:

- Temporal decorrelation
- Geometric decorrelation
- Volume scattering
- Processing error

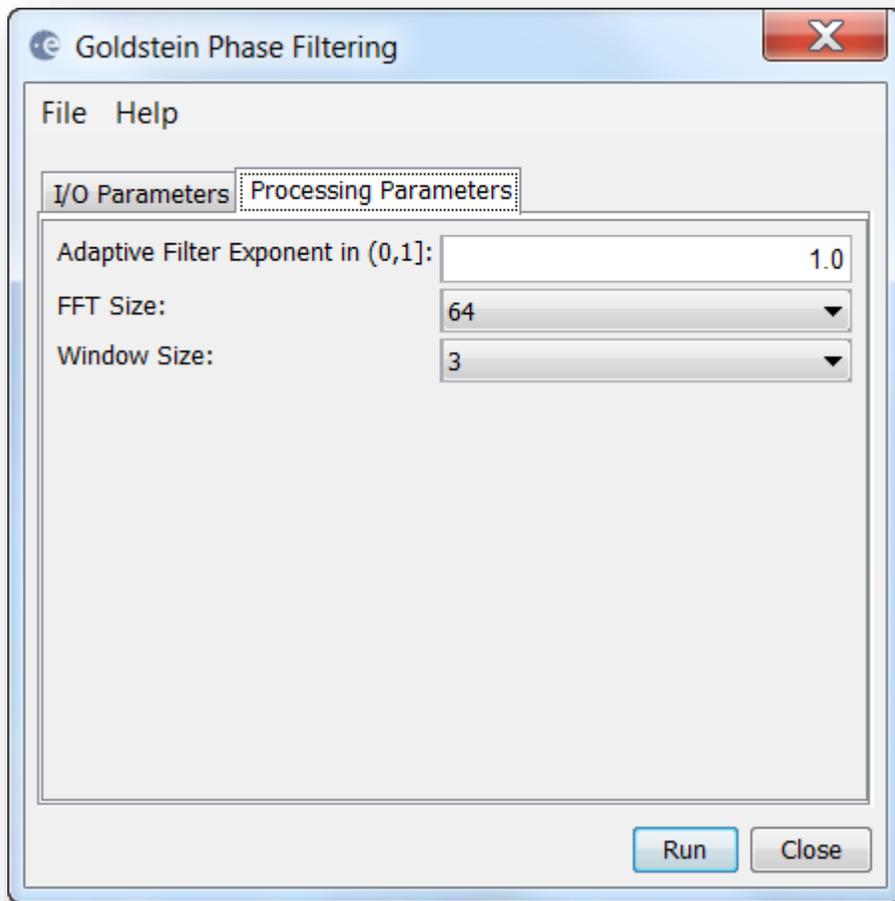
Where there is loss of coherence, the interference pattern is lost.

To be able to properly unwrap the phase, the signal-to-noise ratio needs to be increased by filtering the phase.

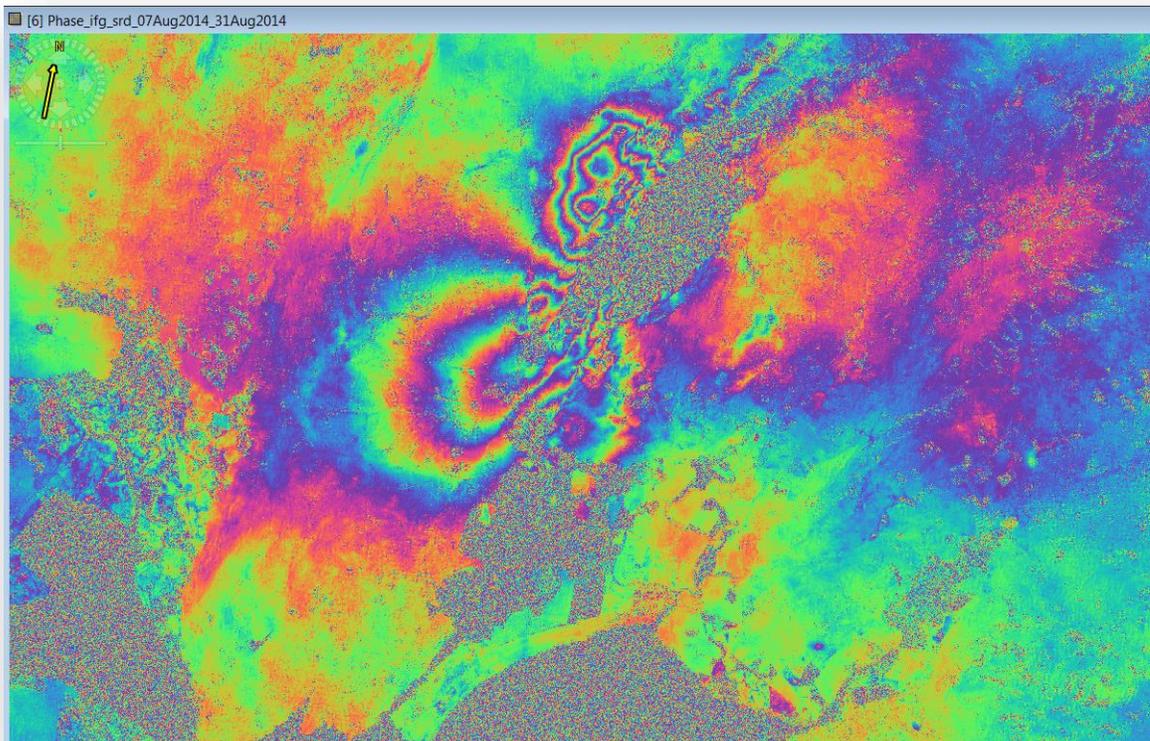
**Step 8 - Phase Filtering:** Select the Interferogram product and go to the **InSAR Tools** menu. Select **Goldstein Phase Filtering**.



Select Phase Filtering



**Phase Filtering Dialog**



**Filtered Phase Band**

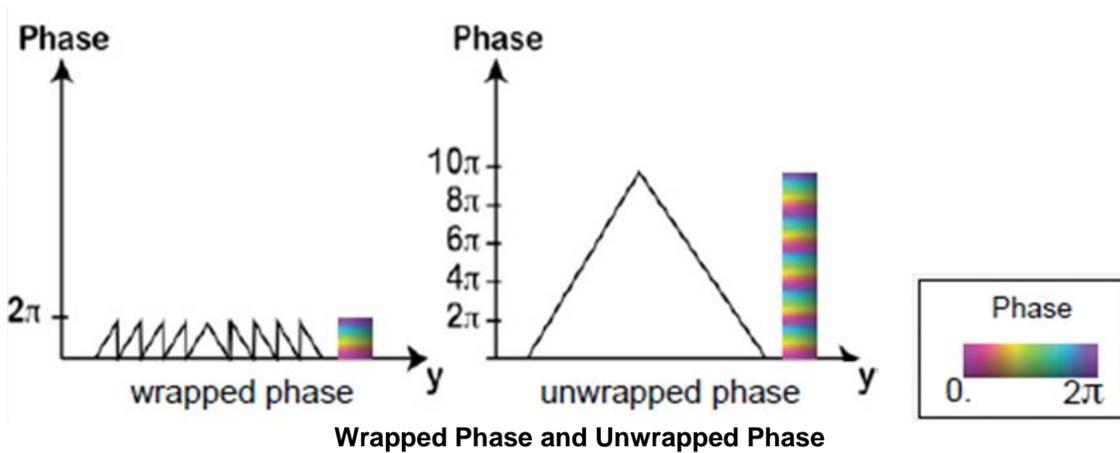
### ***Phase Unwrapping***

In the interferogram, the interferometric phase is ambiguous and only known within  $2\pi$ . In order to be able to relate the interferometric phase to the topographic height, the phase must first be unwrapped.

The altitude of ambiguity  $h_a$  is defined as the altitude difference that generates an interferometric phase change of  $2\pi$  after interferogram flattening.

$$h_a = \frac{\lambda R \sin \theta}{2B_n}$$

Phase unwrapping solves this ambiguity by integrating phase difference between neighbouring pixels. After deleting any integer number of altitudes of ambiguity (equivalent to an integer number of  $2\pi$  phase cycles), the phase variation between two points on the flattened interferogram provides a measurement of the actual altitude variation.

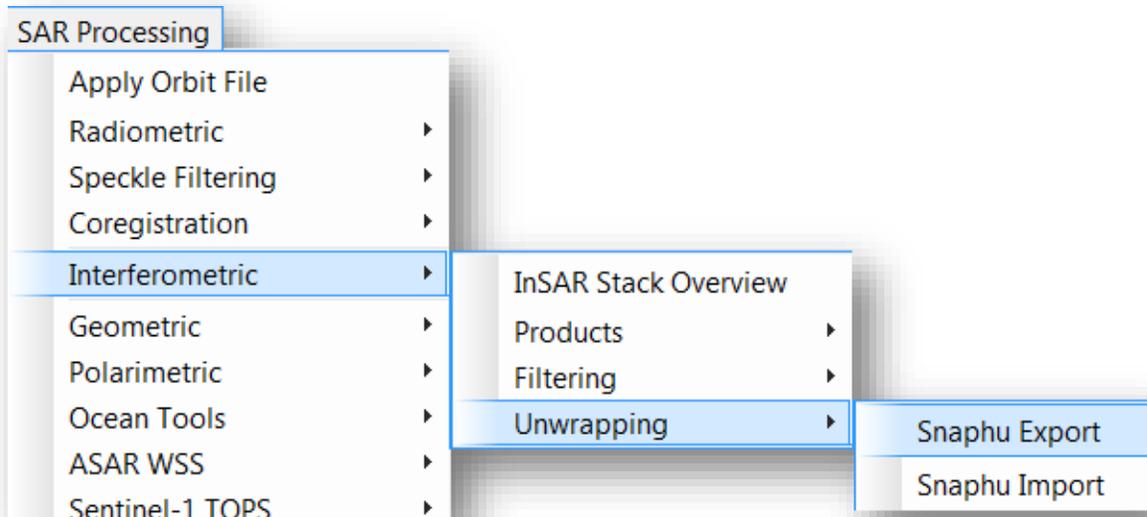


For optimal unwrapping results it is recommended to multi-look (i.e., square) and phase-filter (i.e., increase signal-to-noise and smooth) the interferogram.

The quality and reliability of unwrapped results very much depends on the input coherence. Reliable results can only be expected in areas with high coherence.

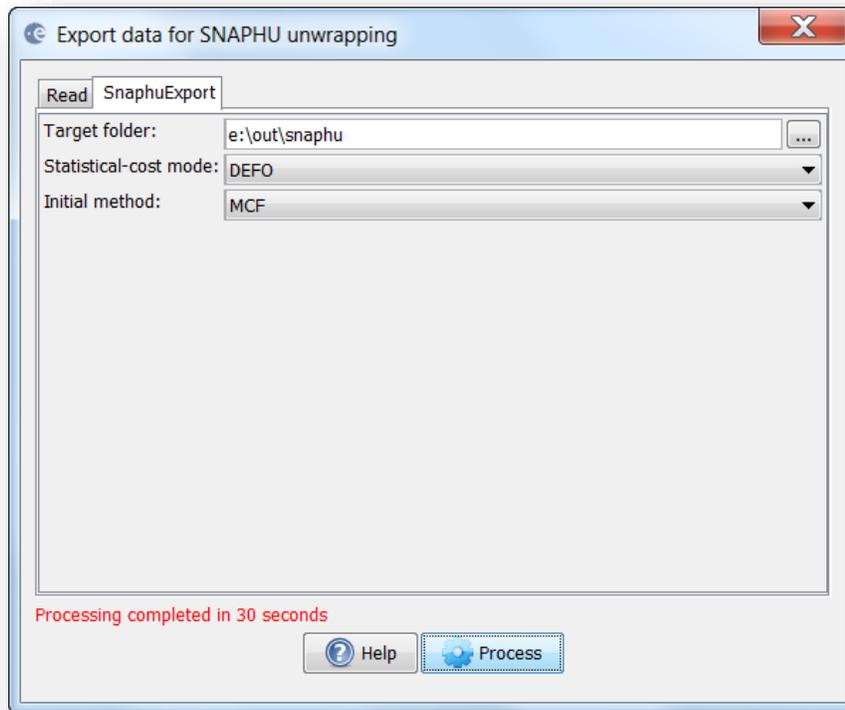
Unwrapped results should be interpreted as a relative height/displacement between two pixels. To obtain absolute estimates, a tie point can be used in the unwrapped phase to height operation.

**Step 8 – Export to Snaphu:** Export the filtered flattened interferogram to SNAPHU.



**Export to Snaphu**

Select TOPO mode for DEM generation.  
 Select DEFO for deformation mapping.



### SNAPHU Export

## ***Unwrapping with SNAPHU***

SNAPHU is a statistical-cost network-flow algorithm for phase unwrapping developed at Stanford University by Curtis Chen and Howard Zebker.

[http://nova.stanford.edu/sar\\_group/snaphu/](http://nova.stanford.edu/sar_group/snaphu/)

Snaphu is available for Linux only. Linux users simply need to install the software package by

**apt-get install snaphu**

Windows users can download a Linux VMWare virtual machine and use it to unwrap the phase.

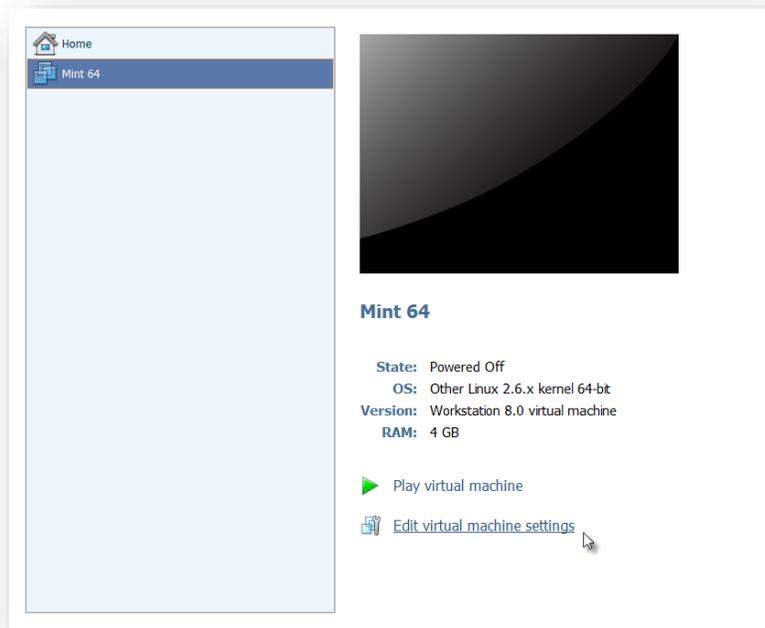
[http://sourceforge.net/projects/s1tbx/files/snaphu\\_vm/SAR%20Mint%2064.zip/download](http://sourceforge.net/projects/s1tbx/files/snaphu_vm/SAR%20Mint%2064.zip/download)

The free VMWare Workstation Player can be downloaded from

<https://my.vmware.com/web/vmware/downloads>

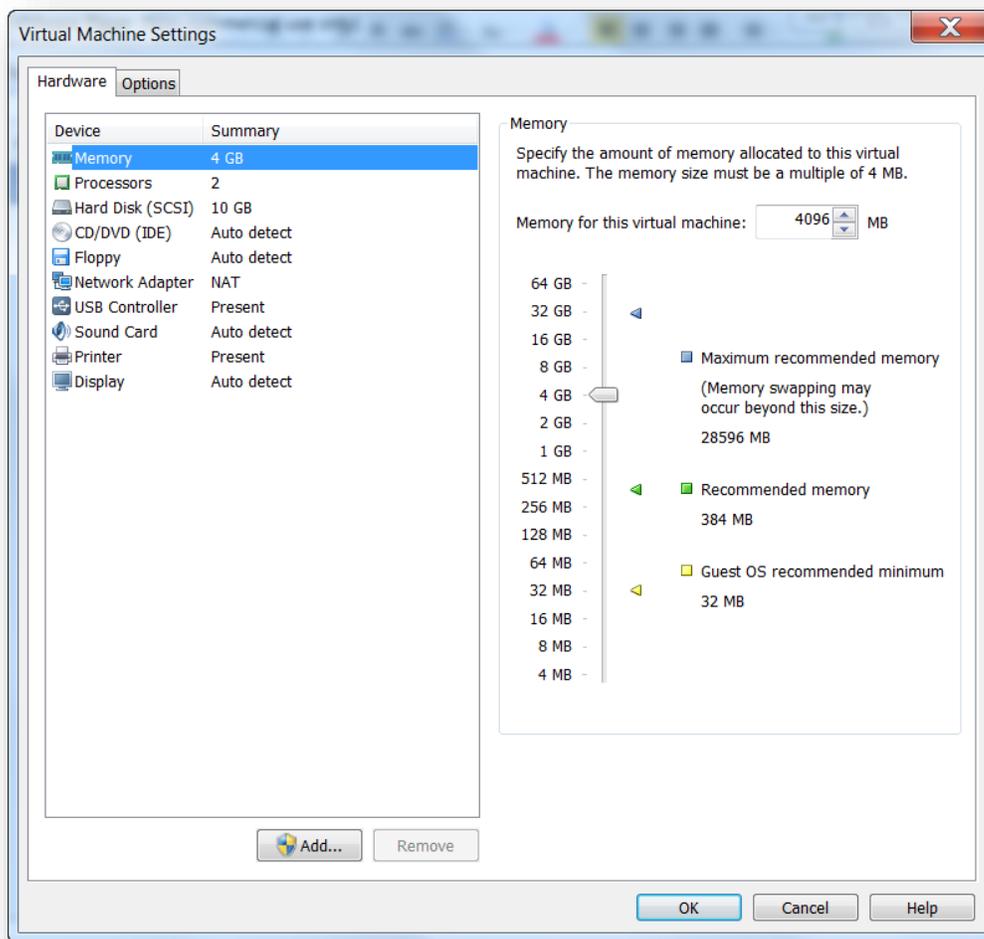
Open the VMWare player and browse for the virtual machine.

Edit the virtual machine settings to increase the memory and setup a shared folder between Linux and Windows.



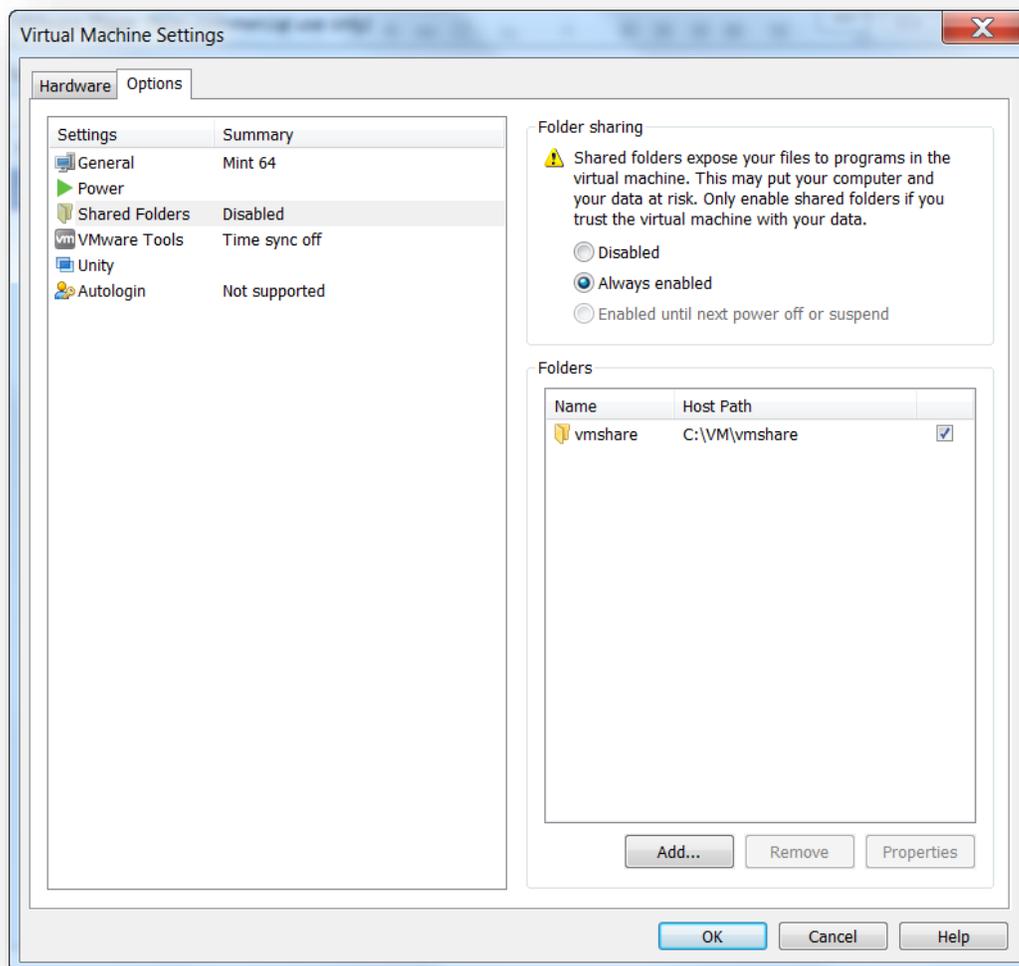
### Edit Virtual Machine Settings

Increase the memory to suit your computer. Depending on the size of your images, you may need at least 8GB.



### Increase Memory

Under the **options** tabs, add a shared folder. Select '**Always Enable**'.



### Enable a Shared Folder

Save the wrapped phase output into the shared folder

After starting the Linux virtual machine, open a command terminal and you should see the data in **/mnt/hgfs/**

Start the virtual machine and login with:

Login: sar

Password: sar01

Go to the data folder in **/mnt/hgfs/** and open the **snaphu.conf** file.

```
cd /mnt/hgfs/vmshare/data/target_snaphu/  
gedit snaphu.conf &
```

```

snaphu.conf ✕
# CONFIG FOR SNAPHU
# -----
# Created by NEST software on: 11:16:49 20/03/2015
#
# Command to call snaphu:
#
#   snaphu -f snaphu.conf Phase_ifg_srd_07Aug2014_31Aug2014.snaphu.img 18453
#####
# Unwrapping parameters #
#####

STATCOSTMODE   DEFO
INITMETHOD     MCF
VERBOSE        TRUE

```

Copy the snaphu command and paste it into the command terminal and then run it.

**snaphu -f snaphu.conf Phase\_ifg\_srd\_07Aug2014\_31Aug2014.snaphu.img 18453**

SNAPHU uses an iterative optimization procedure; its execution time depends on the difficulty of the interferogram.

Unwrapping can use a lot of memory. If the unwrapping fails due to there being not enough memory, you could create a subset of your area of interest and try with SNAPHU again.

```

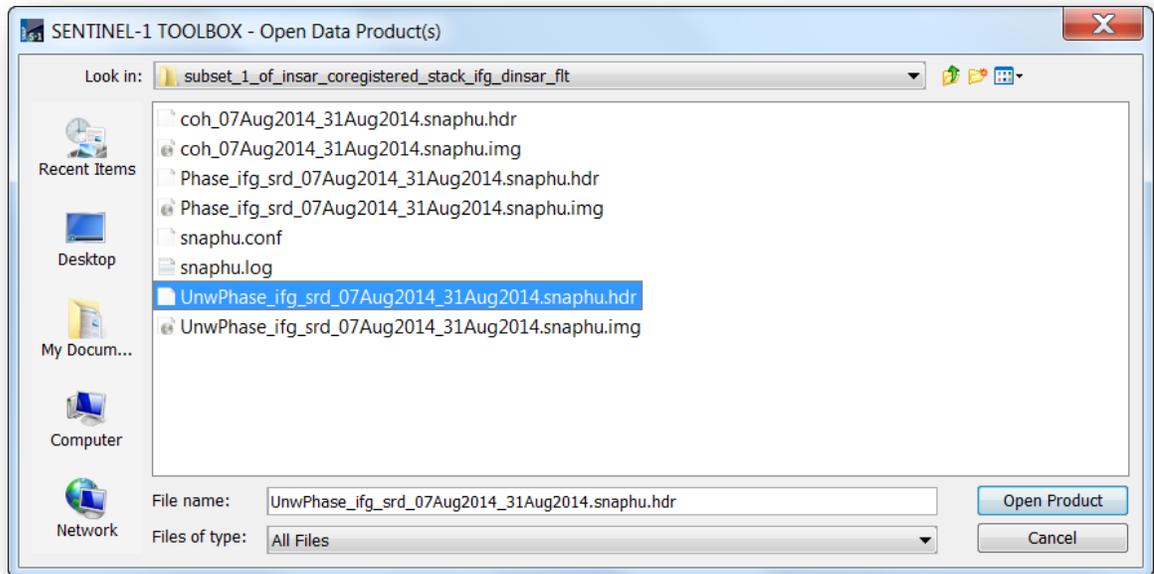
Terminal
Building azimuth cost arrays
Initializing flows with MCF algorithm
Setting up data structures for cs2 MCF solver
Running cs2 MCF solver
Running nonlinear network flow optimizer
Maximum flow on network: 12
Number of nodes in network: 32349493
Flow increment: 1 (Total improvements: 0)
1 incremental cost clipped to avoid overflow (0.000%)
Treesize: 32349493 Pivots: 581827 Improvements: 12560
Maximum flow on network: 3
Flow increment: 2 (Total improvements: 12560)
Treesize: 32349493 Pivots: 20 Improvements: 0
Maximum flow on network: 3
Flow increment: 3 (Total improvements: 12560)
Treesize: 32349493 Pivots: 0 Improvements: 0
Maximum flow on network: 3
Total solution cost: 138196748
Integrating phase
Writing output to file Unw_Phase.img
Program snaphu done
Elapsed processor time: 0:46:23.01
Elapsed wall clock time: 0:46:51

```

**SNAPHU Output**

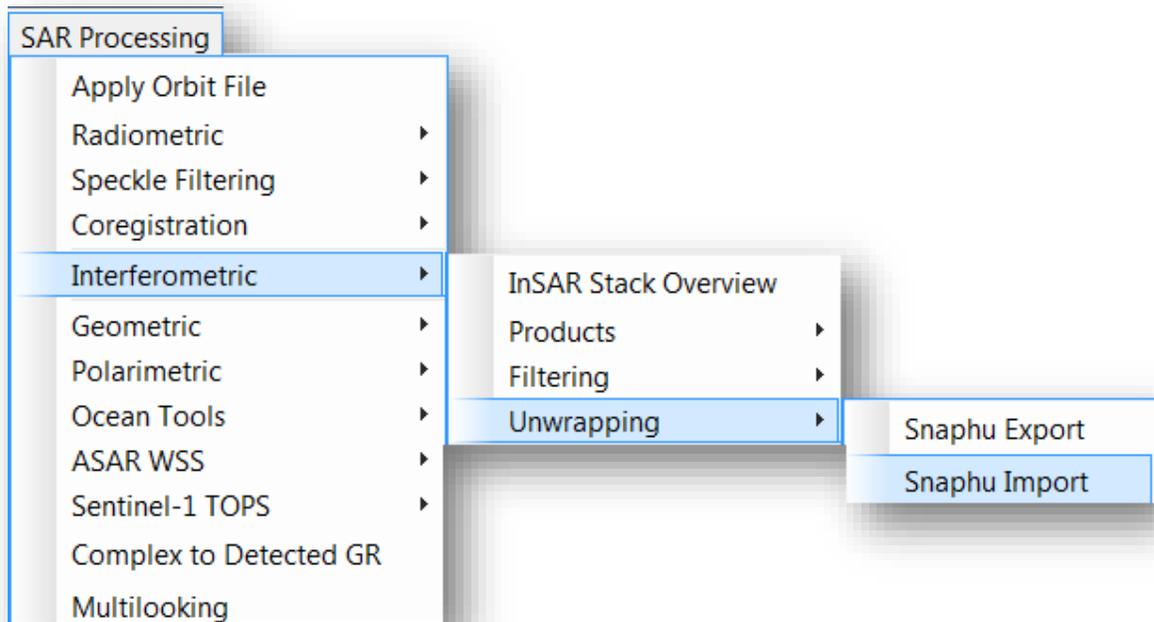
**Step 10 – Import Snaphu Unwrapped Phase:**

Open the Unwrapped phase hdr file



The opened unwrapped phase will not have any metadata or geocodings. We must apply the Snaphu import to join the wrapped phase product with the unwrap phase product.

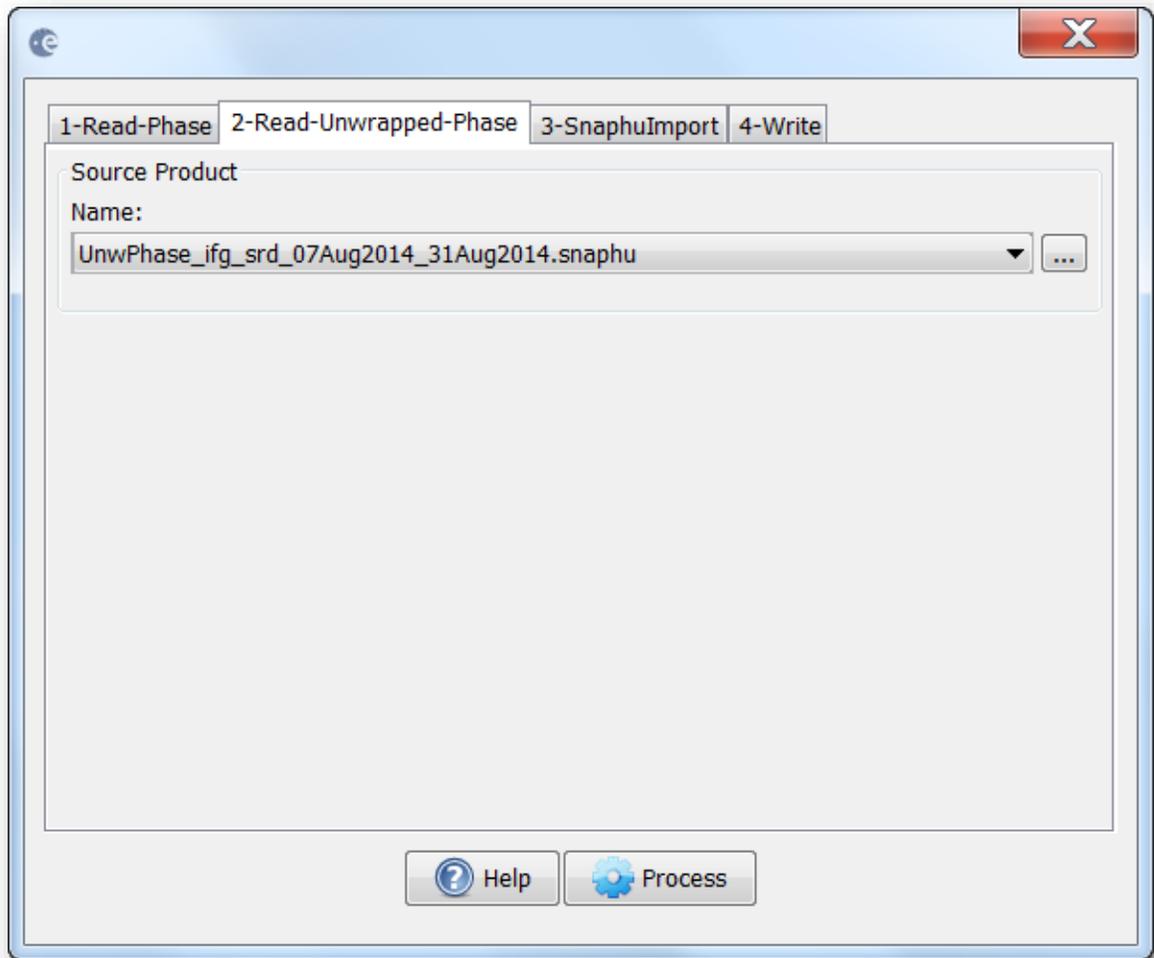
Select **Snaphu Import** from the interferometric menu.



### Snaphu Import

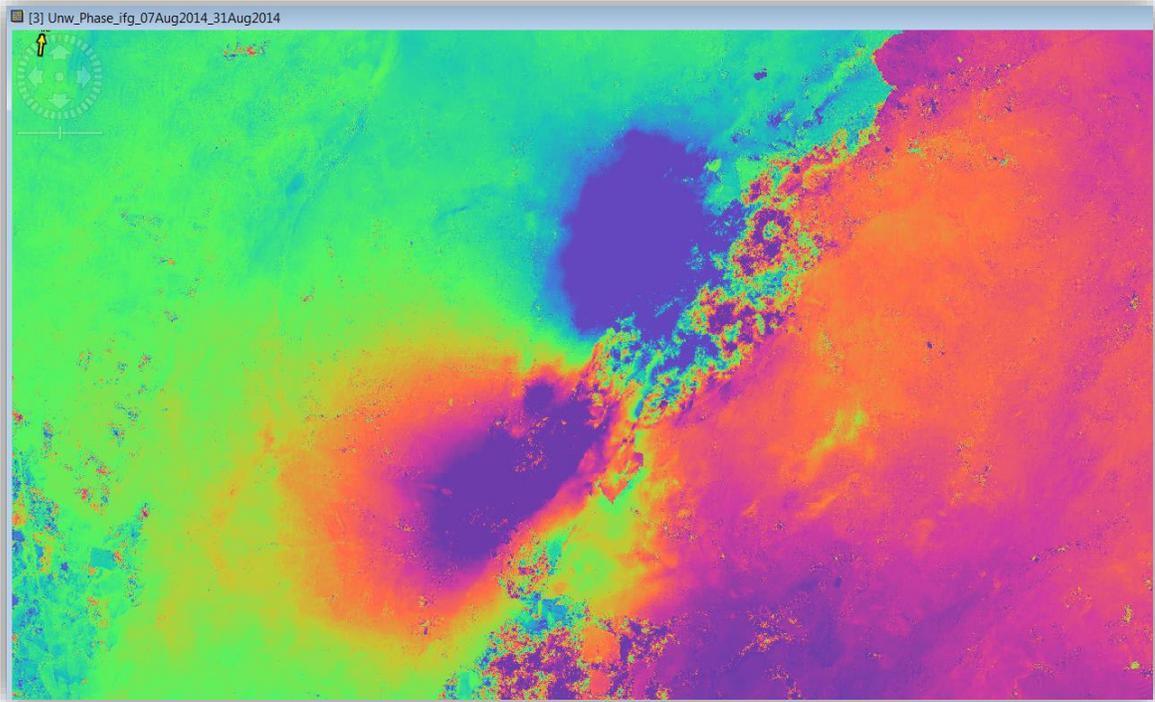
Select the wrapped phase in the **read phase** tab.

Select the unwrapped phase product in the **read unwrapped phase** tab.



**Snaphu Import Dialog**

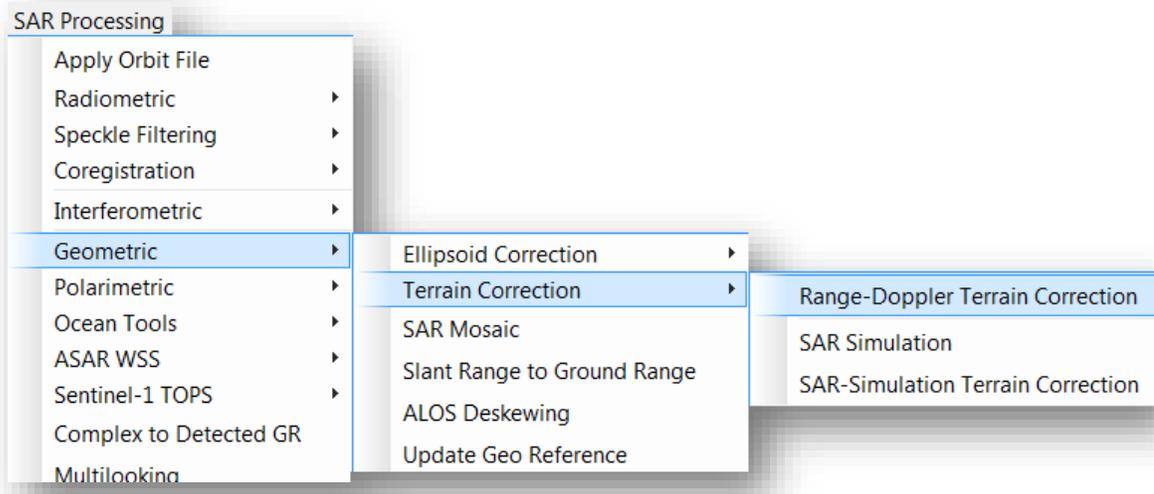
Process and display the output unwrapped phase.



**Unwrapped Phase**

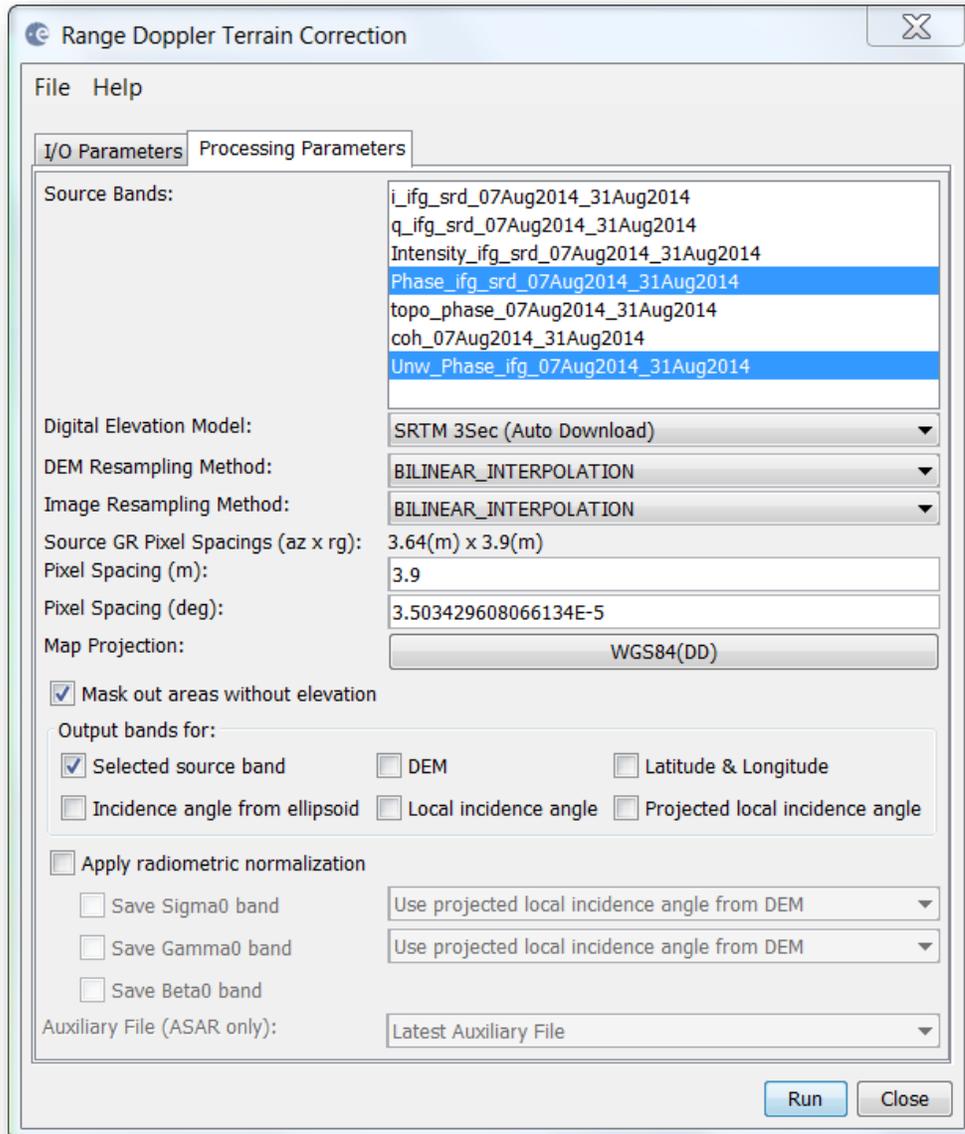
**Step 11 – Geocode the product:**

Apply Terrain Correction to the phase band to geocode the product.

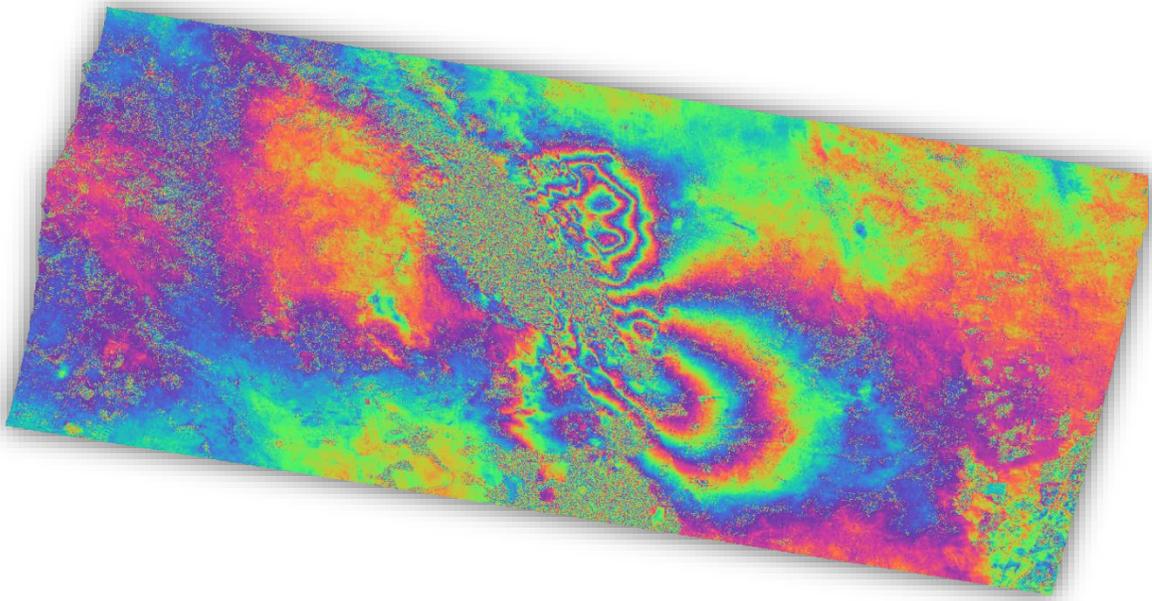


**Apply Range Doppler Terrain Correction**

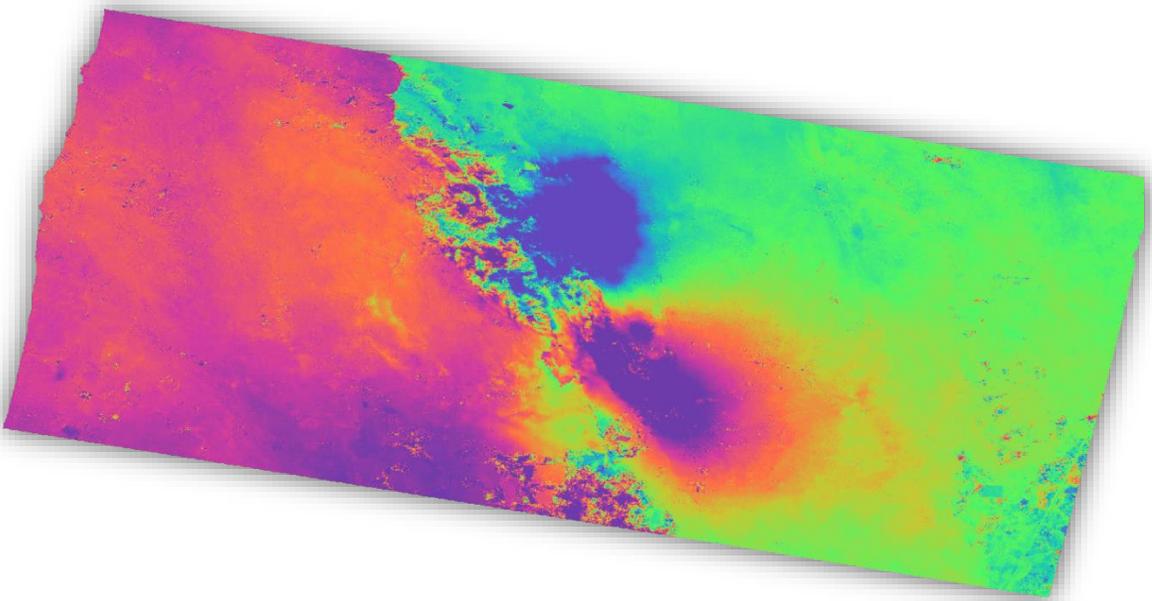
Select the wrapped and unwrapped phase bands.



### Select Phase Bands



**Geocoded Interferogram**



**Geocoded Unwrapped Phase**



For more tutorials visit the Sentinel Toolboxes website

<http://step.esa.int/main/doc/tutorials/>



Send comments to the SNAP Forum

<http://forum.step.esa.int/>